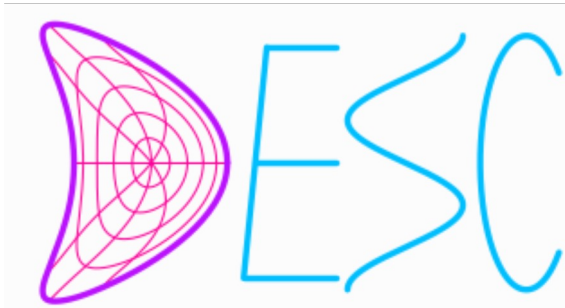


Flexible and Fast Stellarator Optimization with DESC Suite

E. Kolemen, Princeton University

**Daniel Dudt, Rory Conlin, Dario Panici, Rahul Gaur,
Patrick Kim, Kaya Unalmis, Eduardo Rodriguez, Alan Goodman, Aza Jalalvand**



Princeton Plasma Control
control.princeton.edu

Postdoc Opportunity!

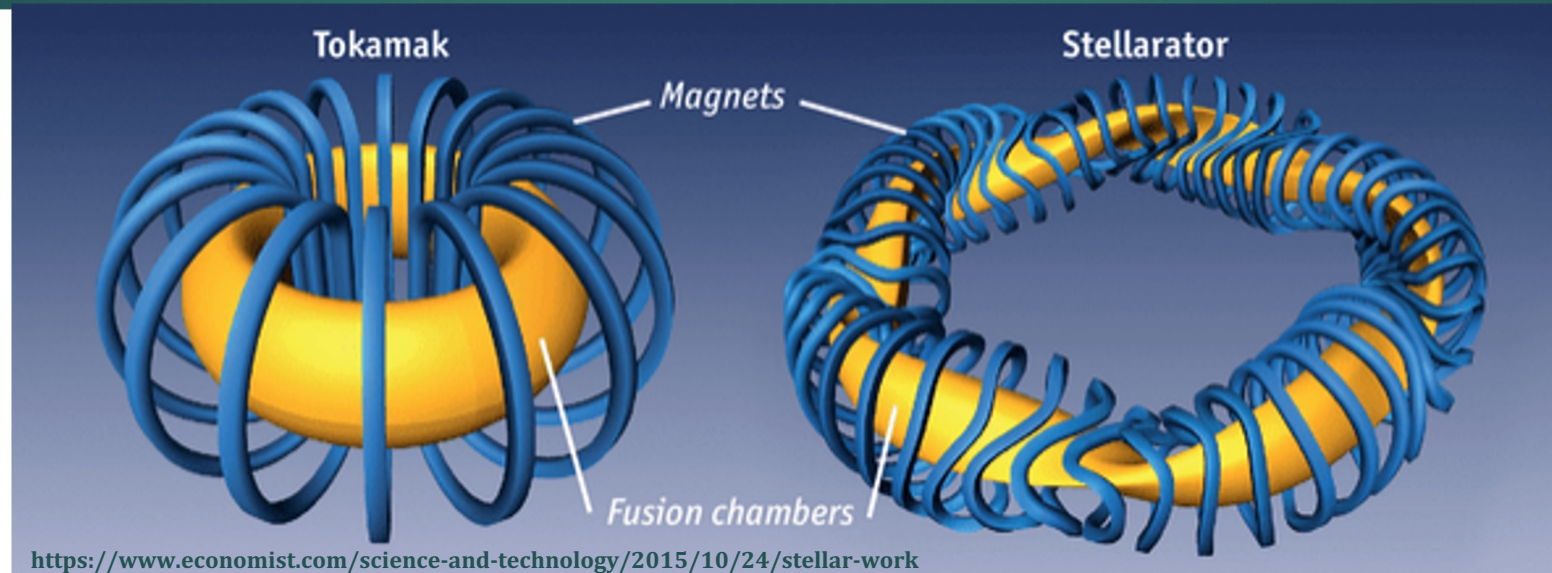
- **Novel Stellarator Designs**
- **Stellarator/Divertor Topology**
- **Plasma SOL**
- **Plasma, Applied Math, CS welcome**
- **Work on DESC**

- **Email: ekolemen@princeton.edu (or talk to me)**

Tokamak Disruption (C-Mod)



Magnetic Confinement Devices: Tokamak vs. Stellarator



Tokamaks

- **Axisymmetric**
 - Simpler geometry
 - Guaranteed particle confinement
- **Requires substantial plasma current**
 - Must be driven
 - NOT steady state!
- **Current leads to instabilities -> disruptions**

Stellarators

- **Inherently 3-D**
 - Complex geometry and coils
 - Confinement not guaranteed
 - certain fields exist which recover this (Quasisymmetry)
 - Larger design space
- **Does not need plasma current**
 - Steady state
 - No disruptions

What is the ideal way to optimize stellarators?

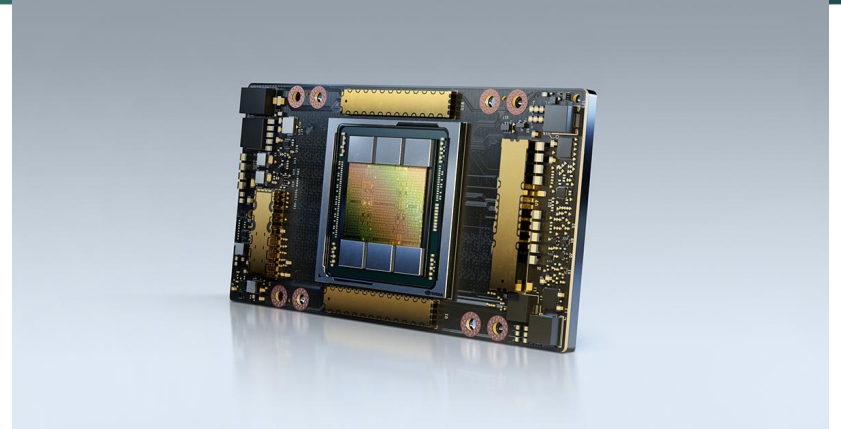
- **Constraints $g(x)$:**
 - MHD equilibrium
 - *Physicist insight: Analytical calculations (e.g. NEA)*
 - *Engineer insight: e.g. $A < 5$, ...*
- **Objectives $f(x)$:**
 - Quasi-symmetry
 - Turbulence
 - ...
- *Physicist/engineer insight: relative importance of $f(x)$*

Then Call A Fast Code

$$\min_x f(x)$$

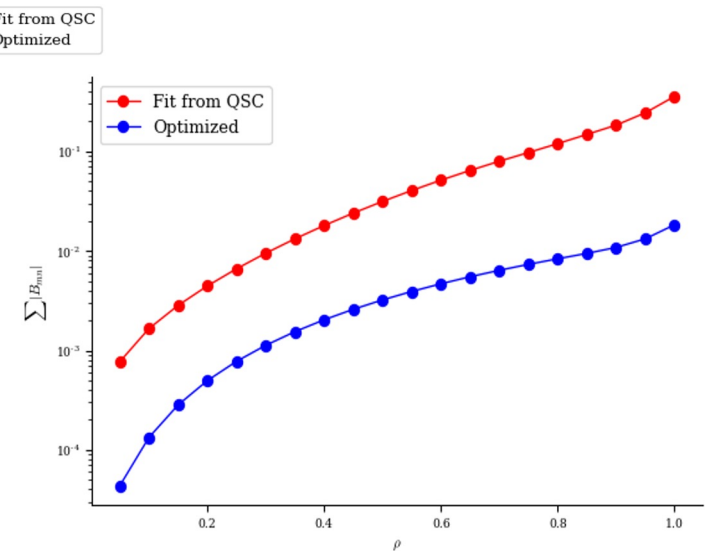
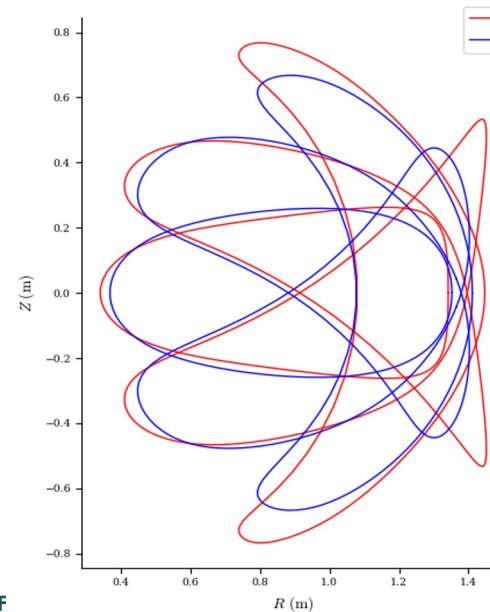
subject to

$$g_{eq}(x) = 0$$
$$g_{ineq}(x) \geq 0$$



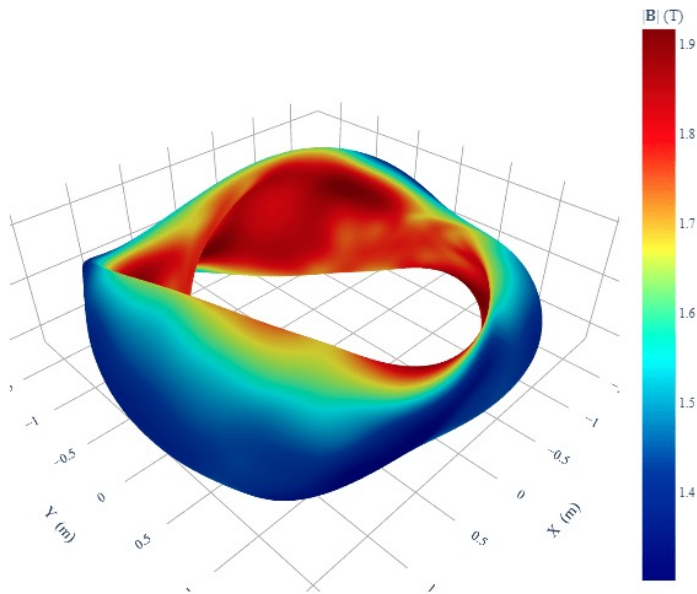
Fast= GPU + Jacobian

**Constraints: $g_{\text{physicist}}$ = Fix NEA
+ g = MHD Eq.
Optimize remaining volume**



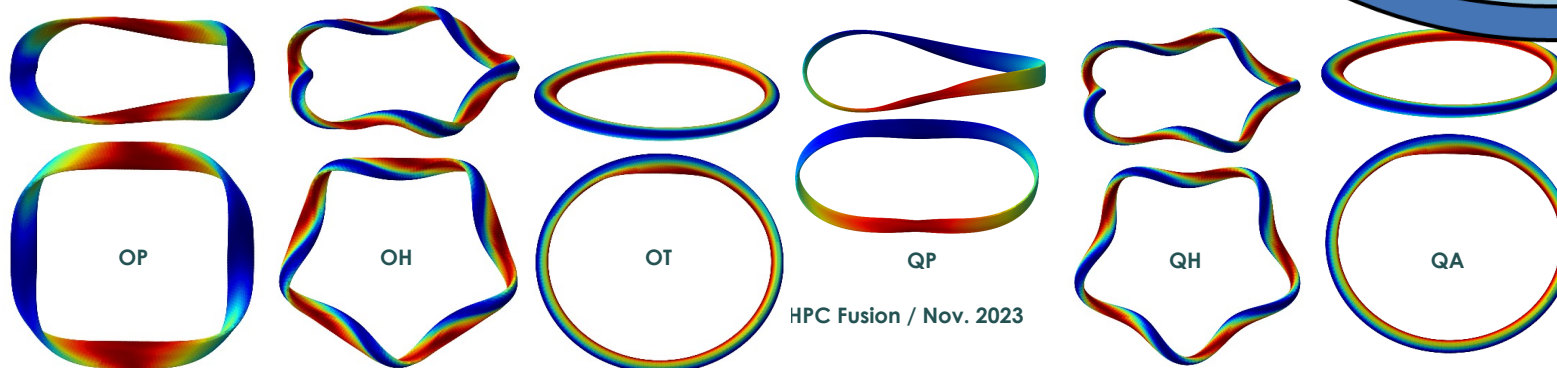
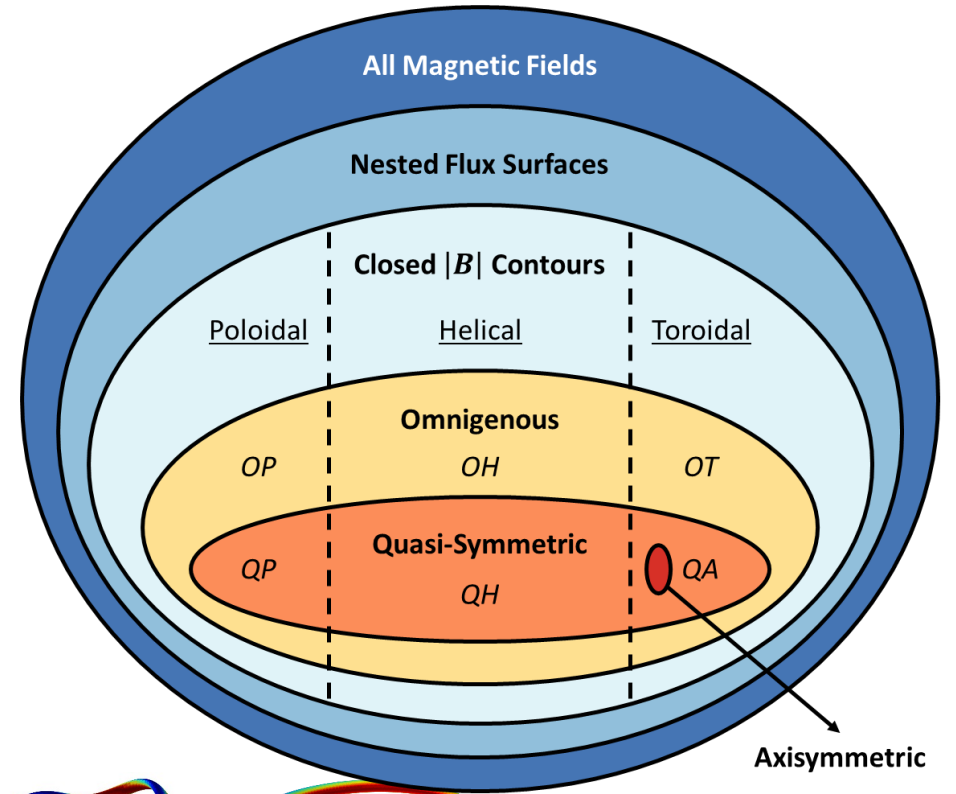
DESC is a new tool for stellarator optimization

Accurate Equilibria

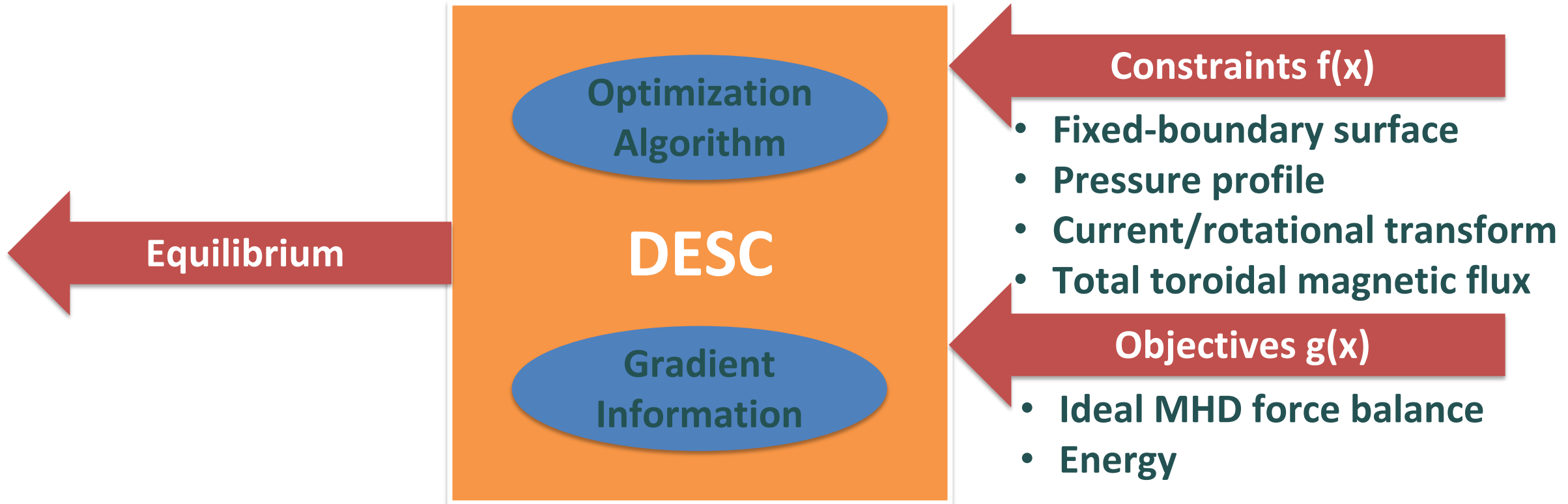


- Stellarator equilibria are complicated
- Design space is much larger than tokamaks

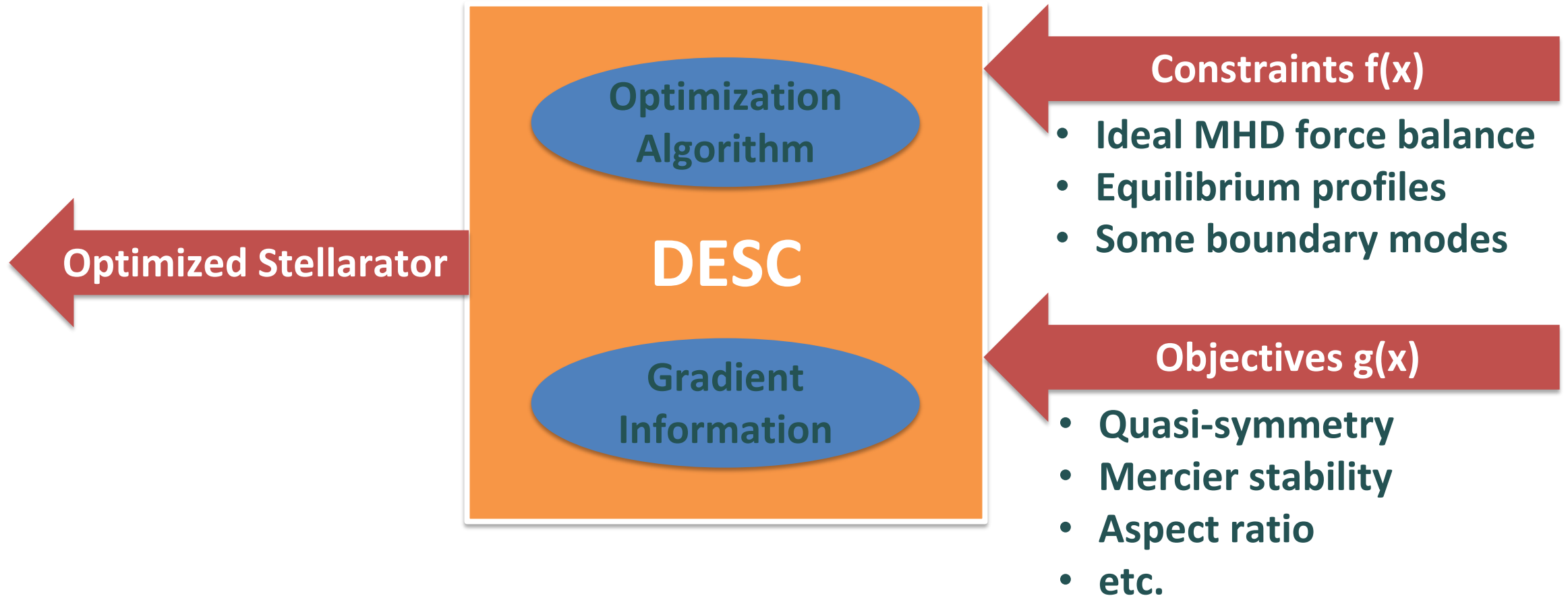
Fast Optimization



A flexible stellarator optimization suite



A flexible stellarator optimization suite



DESC Developed with the following design principles:

1. Simple user interface

- Open-source Python code
- Well documented
- High test coverage
- Easy to install

2. Local error quantification

- Pseudo-spectral (collocation) methods

3. Properly resolve the magnetic axis

- Global basis functions
- Zernike polynomials

4. Exact derivatives of all objectives

- Automatic differentiation (JAX)

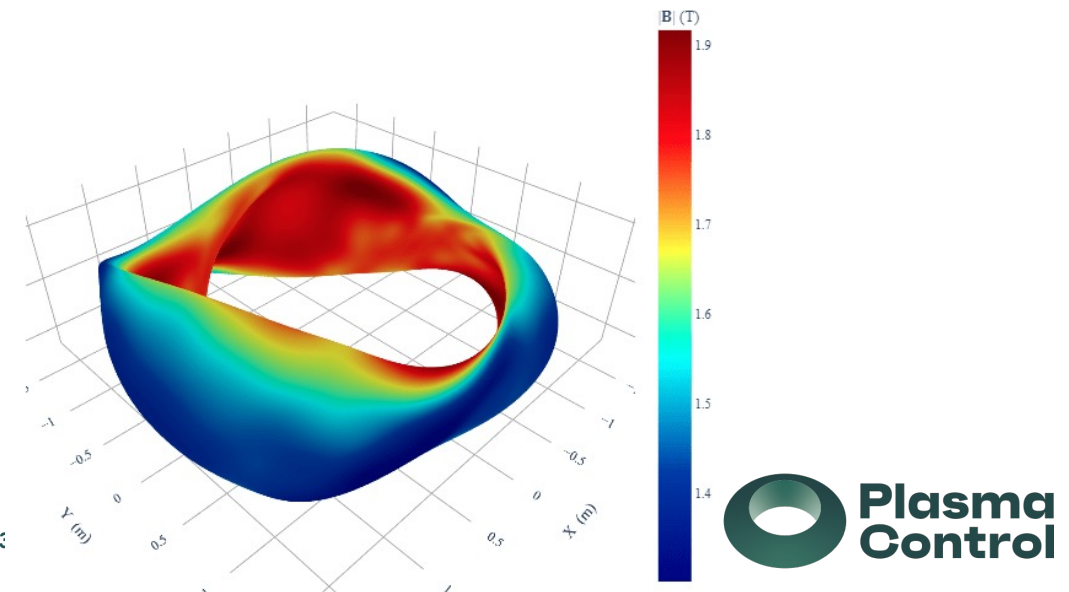
5. Hardware agnostic

- Run on CPUs, GPUs, and TPUs

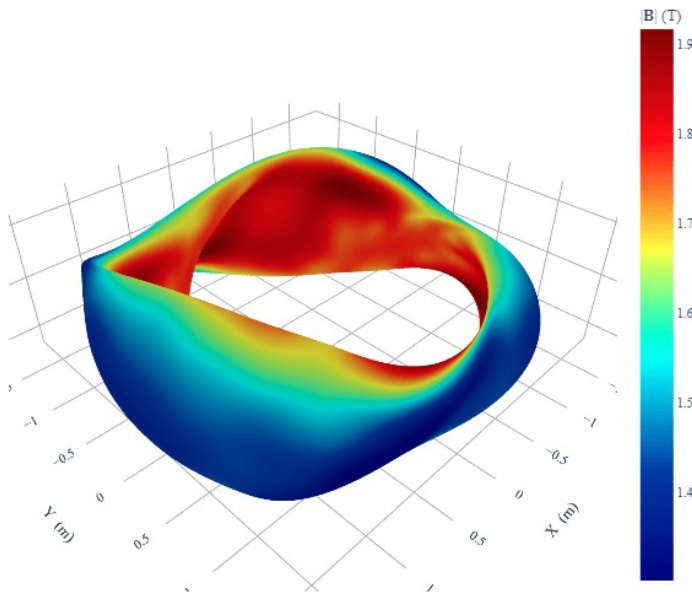
6. Extendable to new applications

- Modular & flexible code structure

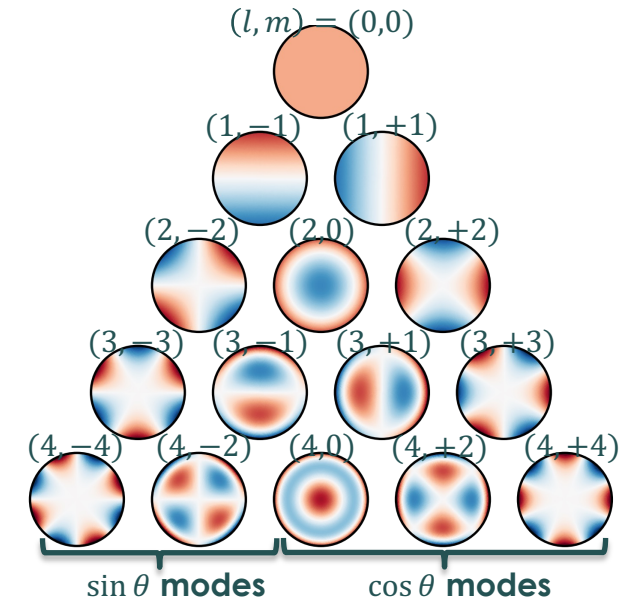
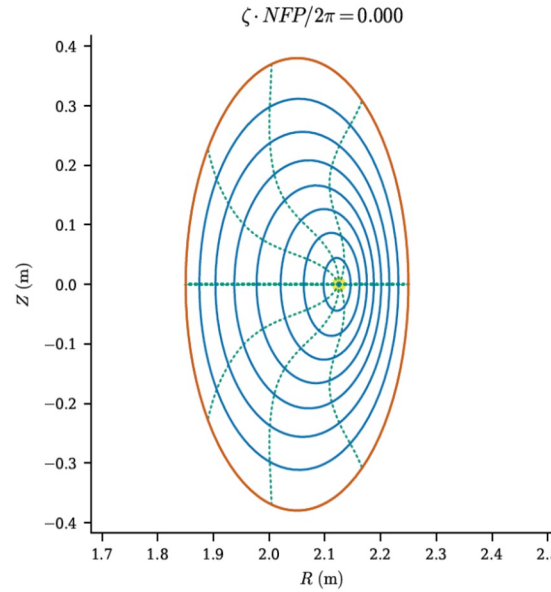
<https://github.com/PlasmaControl/DESC>



Zernike spectral basis better than concentric circle basis



Poincare section



spectral coefficients

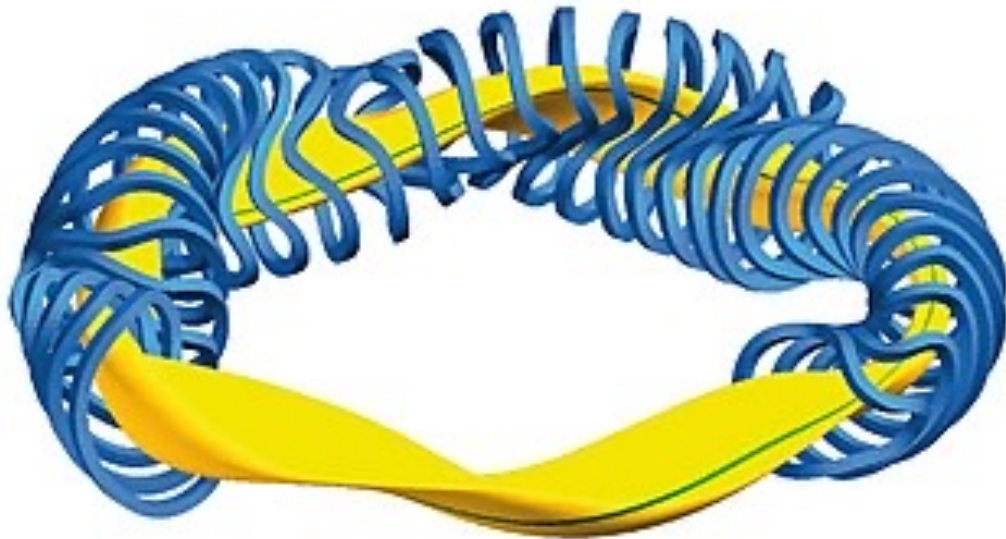
Zernike polynomials

$$X(\rho, \theta, \zeta) = \sum_{lmn} X_{lmn} Z_l^m(\rho, \theta) \mathcal{F}^n(\zeta)$$

Fourier series

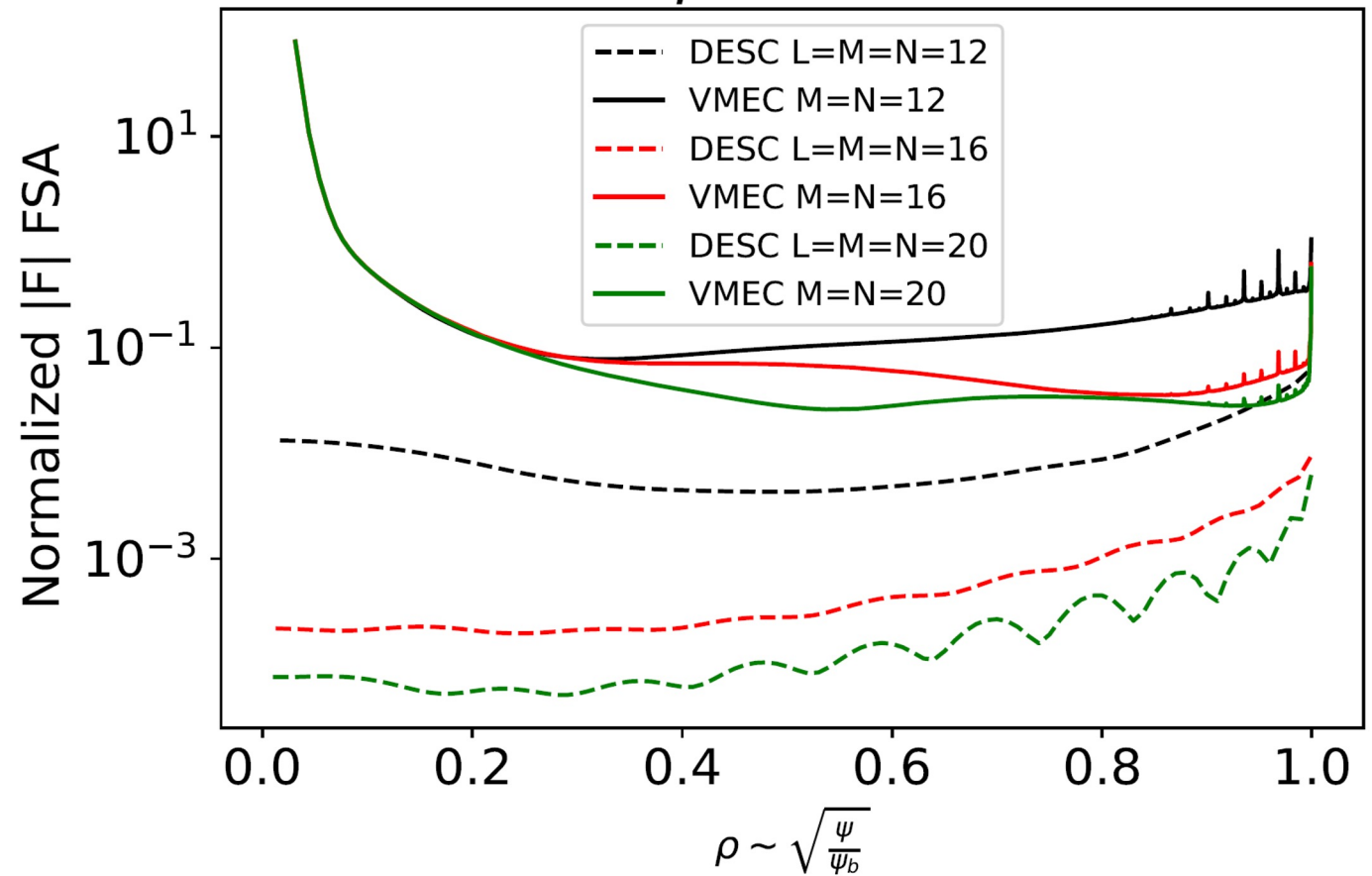
- Satisfies analyticity conditions at the magnetic axis
- Exponential convergence (if solution exists and is smooth)

DESC spectral methods yield more accurate equilibrium solutions

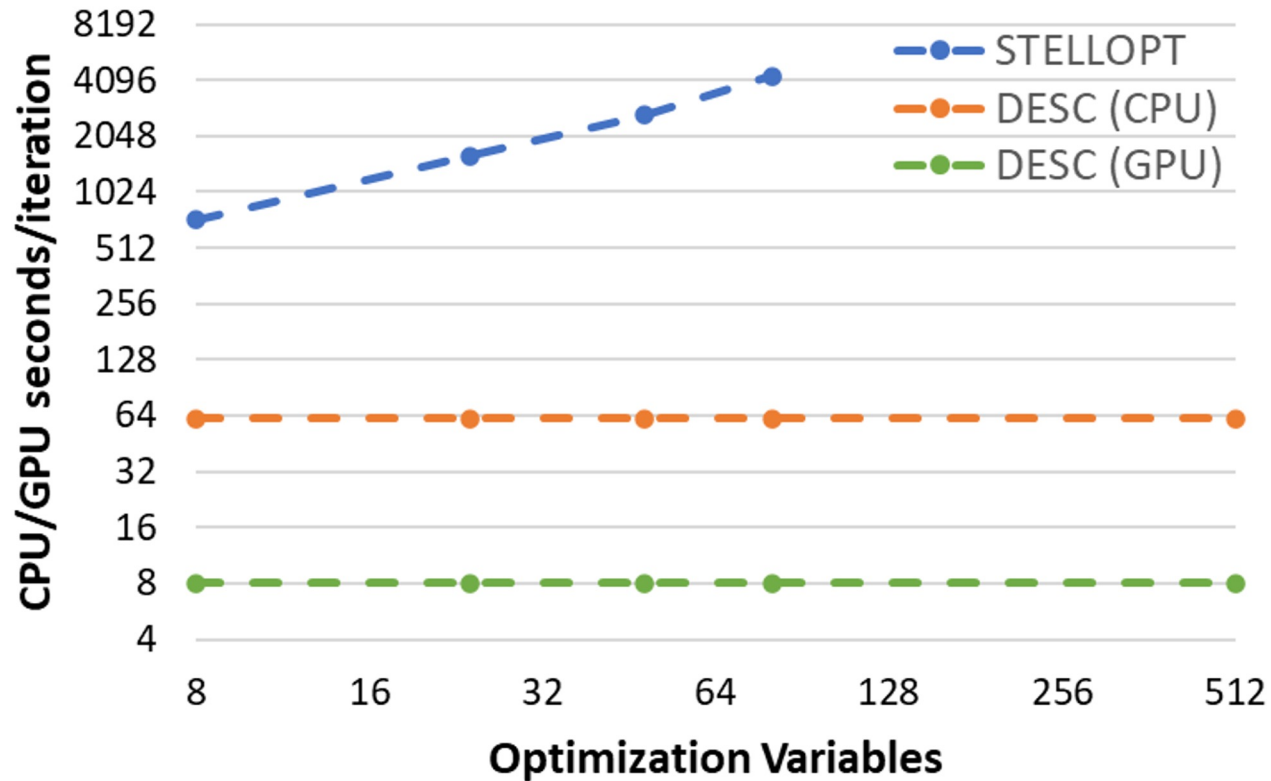


W7-X

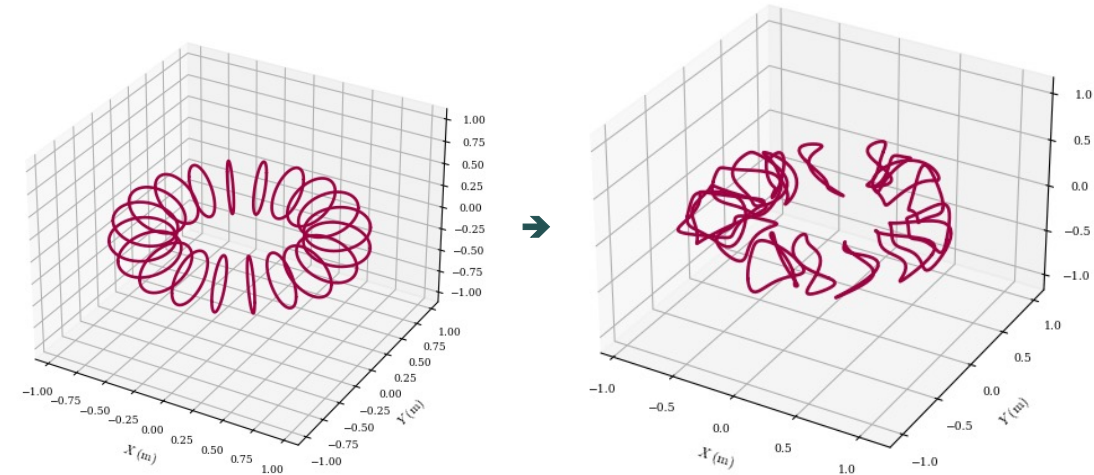
W7-X 2% β FSA Force Error



DESC Allow Much Faster Stellarator Optimization



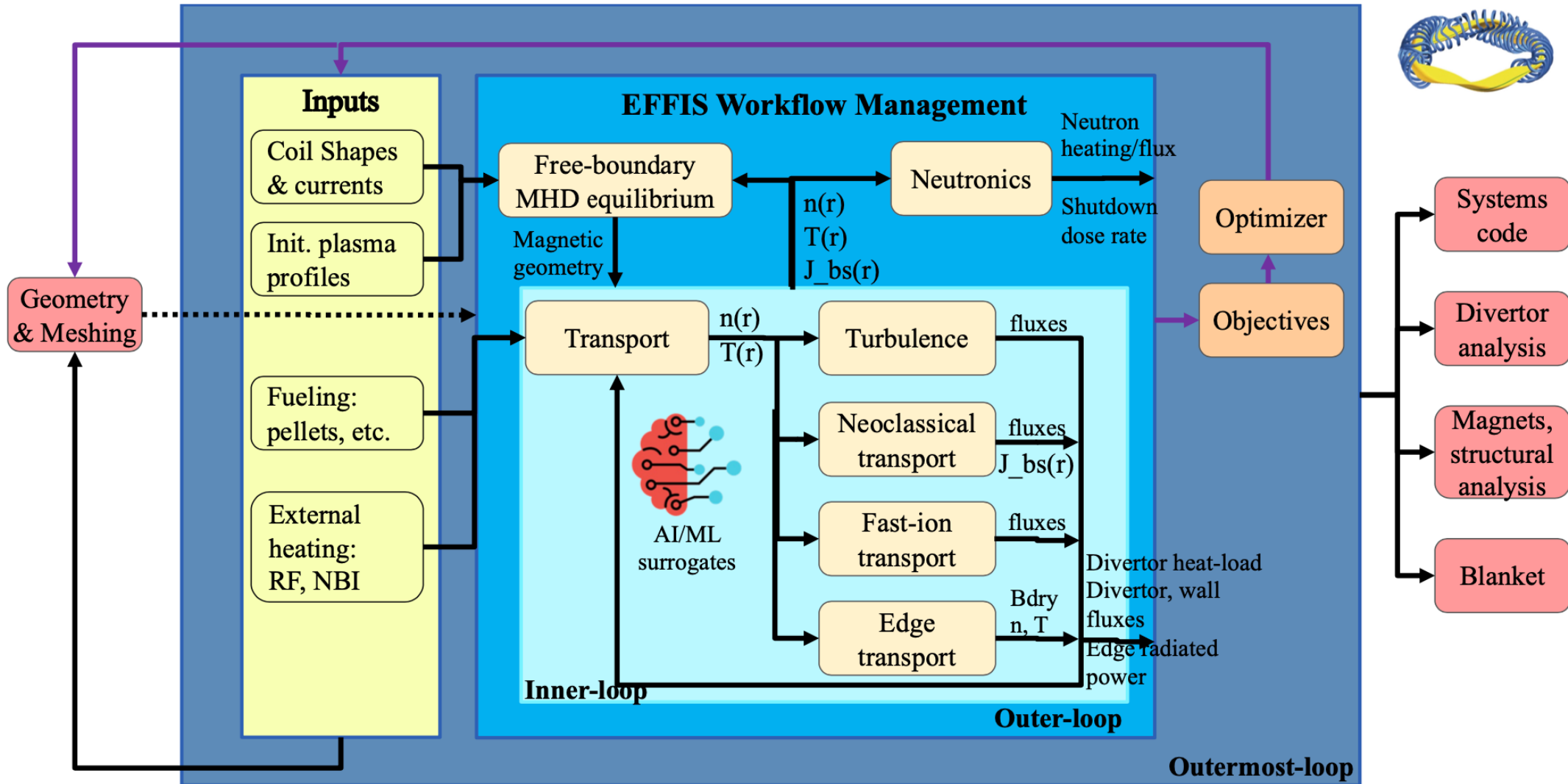
Example Coil Optimization



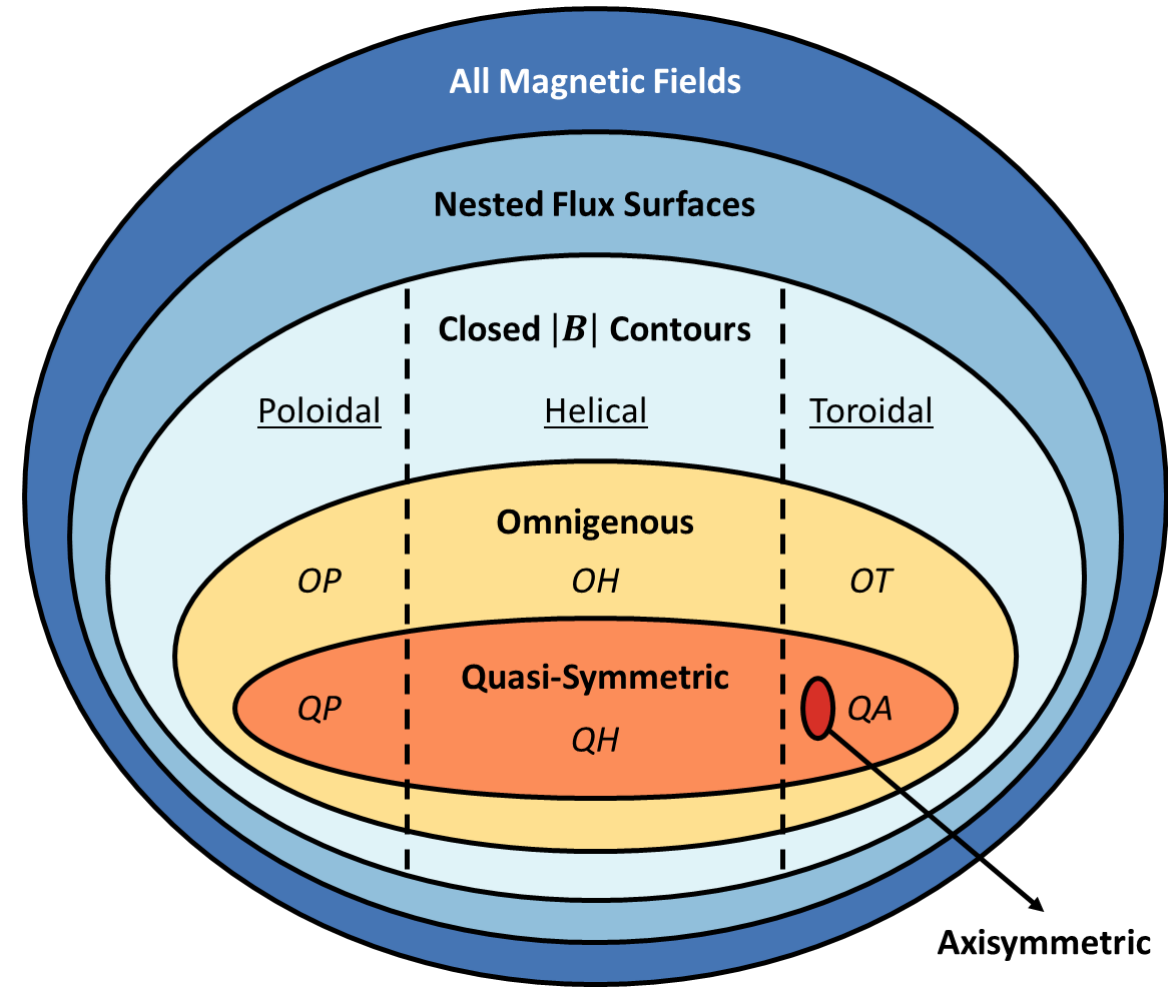
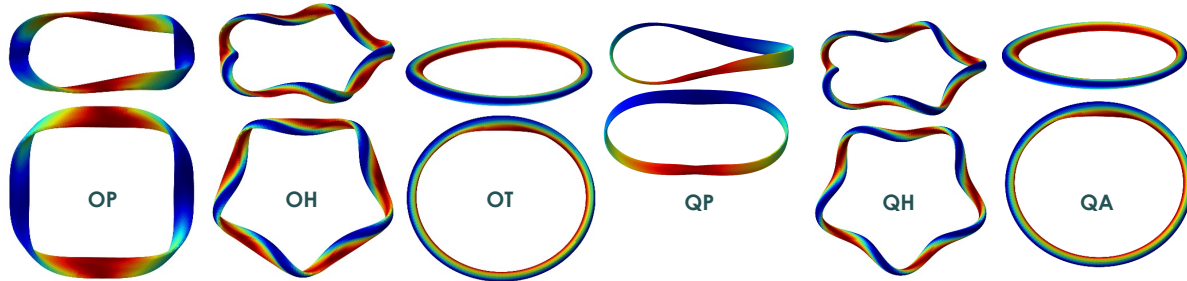
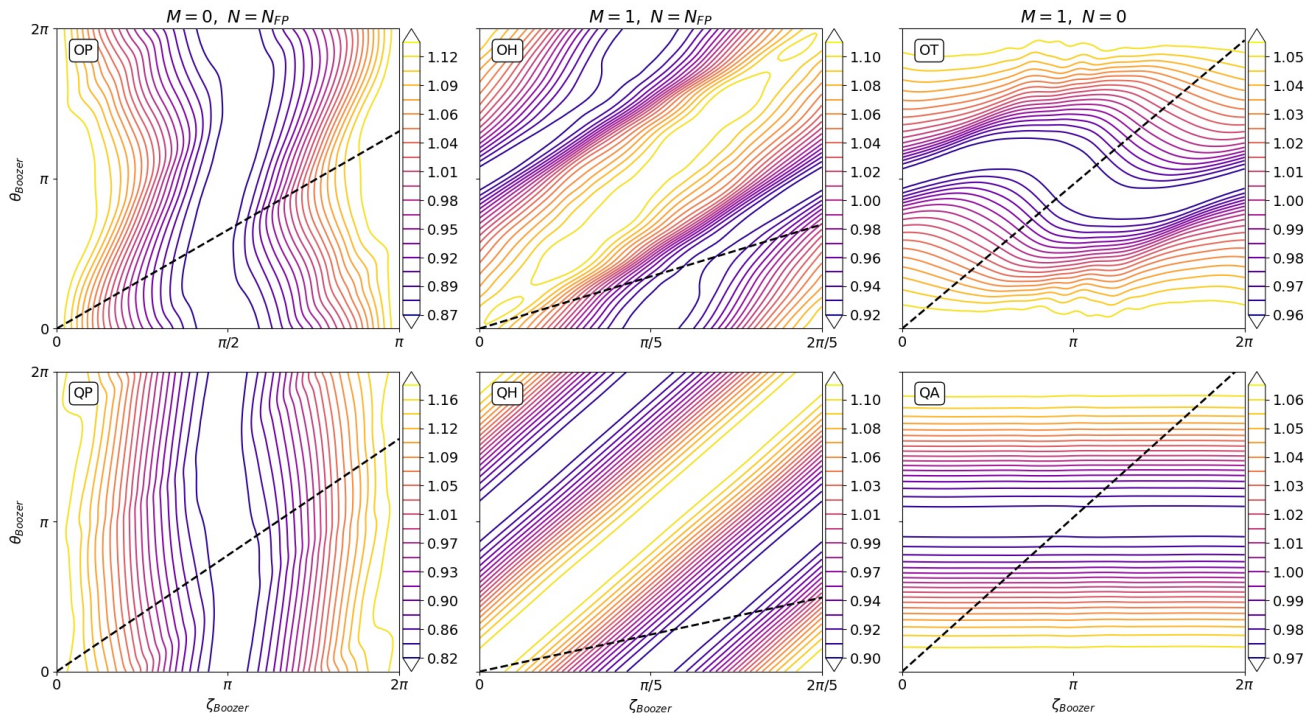
$$W7-X \beta = 2\%; L = 24, M = N = 12$$

Hardware	Run Time
Intel Cascade Lake CPU	48 min
NVIDIA A100 GPU	20 min

Next up: End-To-End Stellarator Optimization with AutoDiff



Main Result: General Omnigenity Easily Obtained with DESC



Current optimization methods: Sum of Squares

Combine equality + inequality constraints

$$\min_x f(x) + w_1 [g(x)]^2$$

Choose small weight for inequality constraints to enforce “approximately”

Choose large weight for equality constraints to penalize a lot

Limitations:

- Hard to guess a-priori what weights should be
- Even small weights for “inequality” constraints can overly penalize things we don’t care about

Better methods: Augmented Lagrangian

- **Combination of traditional Lagrangian + quadratic penalty**

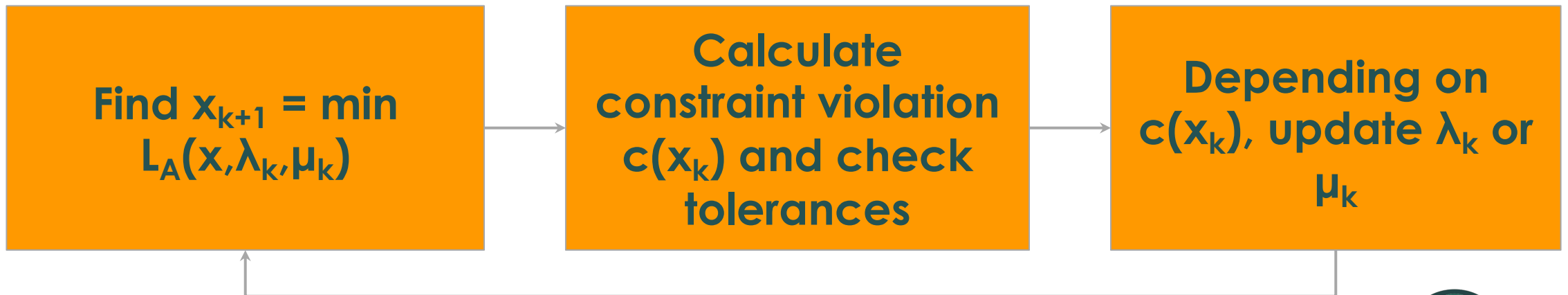
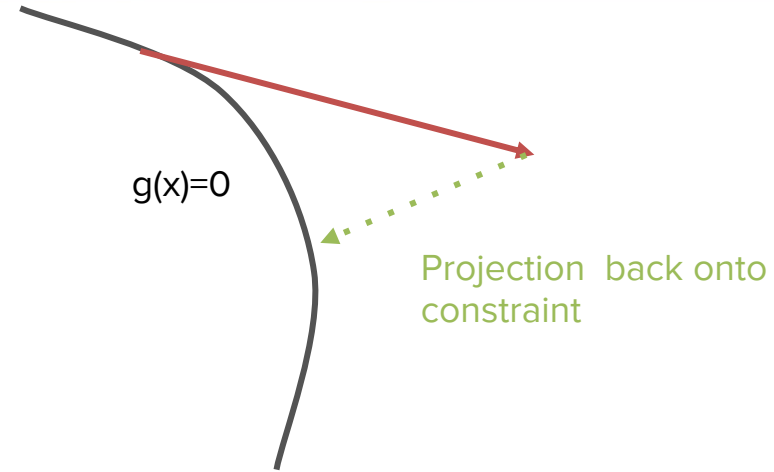
$$\mathcal{L}(x, \lambda, \mu) = f(x) + \lambda^T \mathbf{g}(x) + \mu g^2(x)$$

- **Doesn't introduce any non-smooth terms**
- **“Exact” method - doesn't need $\mu \rightarrow \infty$**
- **Solve sequence of subproblems for increasing μ, λ**
- **Provides estimate of true Lagrange multipliers - useful information about trade-offs**

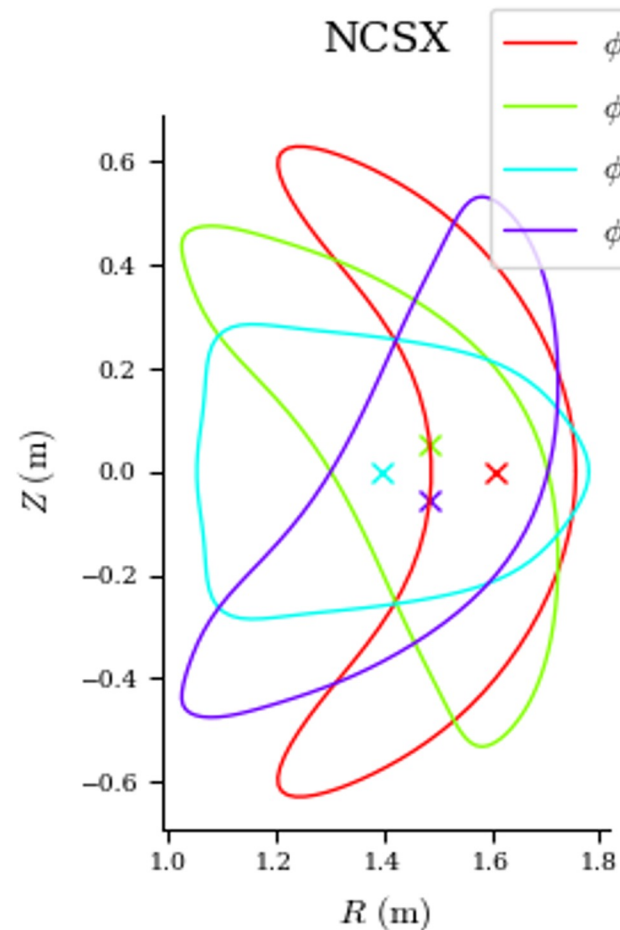
- **Open source packages available (LANCELOT, NLOpt, etc). Also python/JAX version implemented in DESC**

DESC Allow Combined Constraints + Optimization

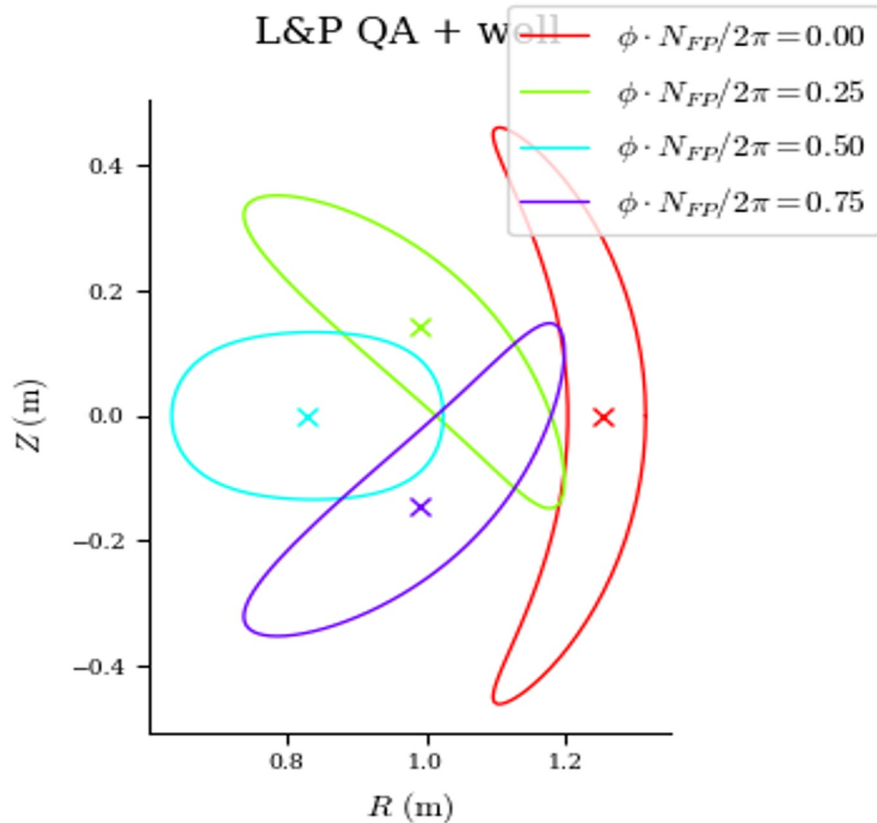
```
1 while optimization stopping criteria are not met
  :
2
3   perturb equilibrium solution to improve
   objective
4
5   while equilibrium stopping criteria are not
   met:
6
7     solve equilibrium force balance
```



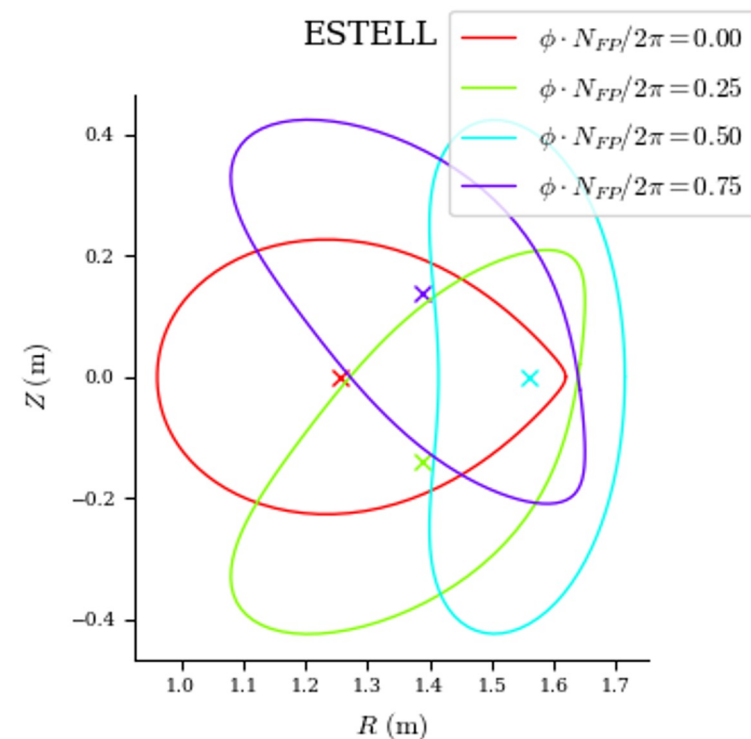
Application of Multi-Objective Optimization in DESC



NCSX + L&P precise
QA: Magnetic well, but
large concave regions



ESTELL: simpler geometry,
but magnetic hill



Application of Multi-Objective Optimization in DESC

Minimize f_{QS} (“two term” metric) subject to

Solve with least squares augmented Lagrangian method in DESC

Force balance: $\mathbf{J} \times \mathbf{B} - \nabla p = 0$

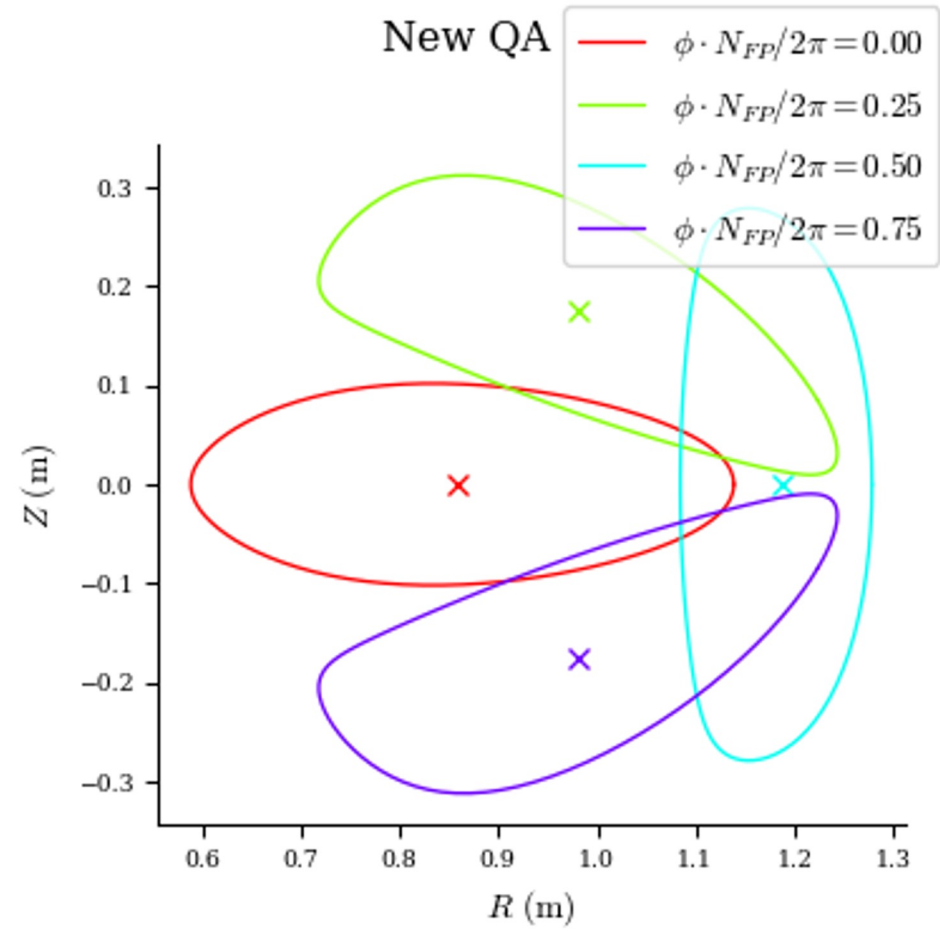
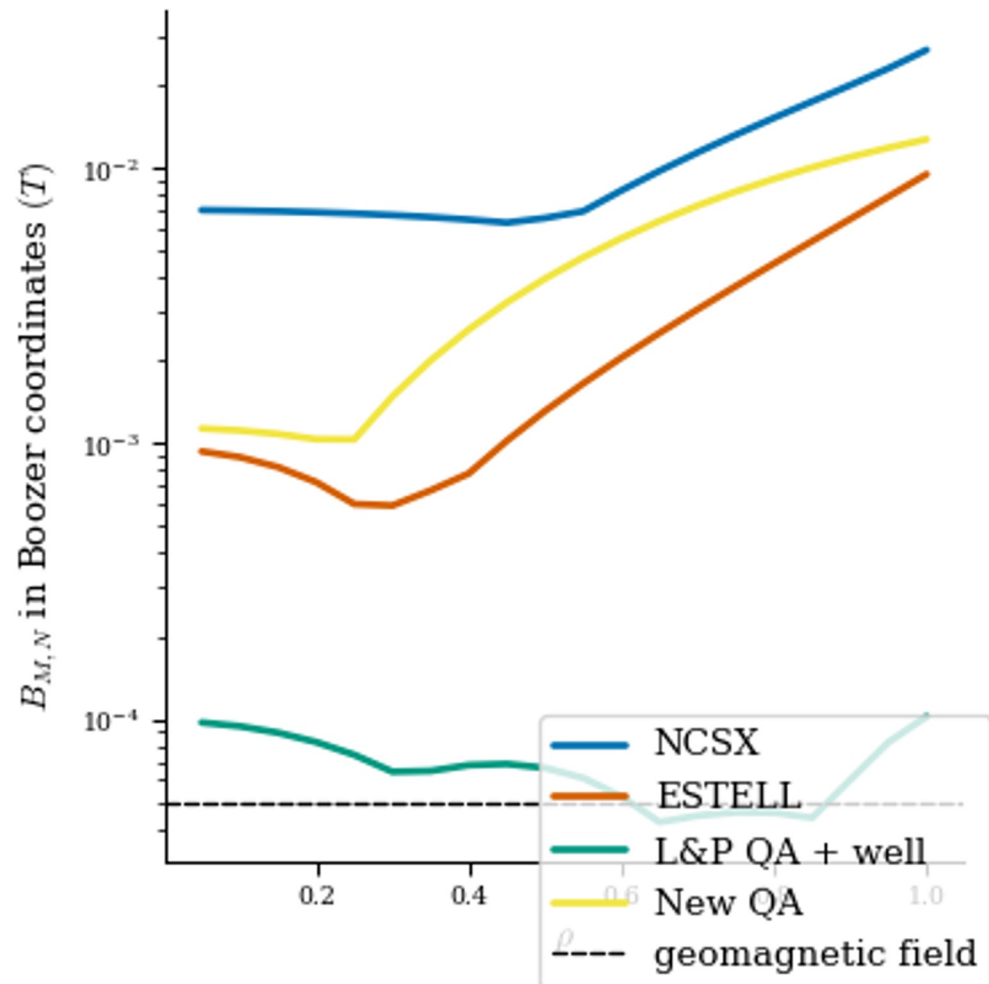
Iota: $0.41 \leq \iota \leq 0.43$

Magnetic Well: $\frac{\partial_\rho V (2\mu_0 \partial_\rho p + \partial_\rho \langle B^2 \rangle)}{V \langle B^2 \rangle} > 0$

Mean Curvature: $H = (\kappa_1 + \kappa_2)/2 < 0$

κ_1, κ_2 eigenvalues of second fundamental form

QS better than NCSX + stability without concavity



Free Boundary DESC

Free boundary constraints:

Flux surface:

$$\mathbf{B} \cdot \mathbf{n} = 0$$

Pressure balance:

$$B_{\text{out}}^2 - B_{\text{in}}^2 - 2\mu_0 p = 0$$

Tangential jump:

$$\mathbf{n} \times [\mathbf{B}_{\text{out}} - \mathbf{B}_{\text{in}}] - \mu_0 \mathbf{K} = 0$$

$\mathbf{B}_{\text{in}} = \mathbf{B}$ from fixed boundary DESC

$$\mathbf{B}_{\text{out}} = \mathbf{B}_{\text{coil}} + \mathbf{B}_{\text{VC}} + \mathbf{B}_{\text{K}}$$

- Parameterize unknown sheet current \mathbf{K} , along with R, Z
- Use high order singular integration scheme (Malhotra et al 2019) to compute virtual casing + sheet current field $\mathbf{B}_{\text{VC}} + \mathbf{B}_{\text{K}}$
- Minimize all 3 equations simultaneously

Free Boundary DESC

Free boundary constraints:

Pressure balance:

$$B_{\text{out}}^2 - B_{\text{in}}^2 - 2\mu_0 p = 0$$

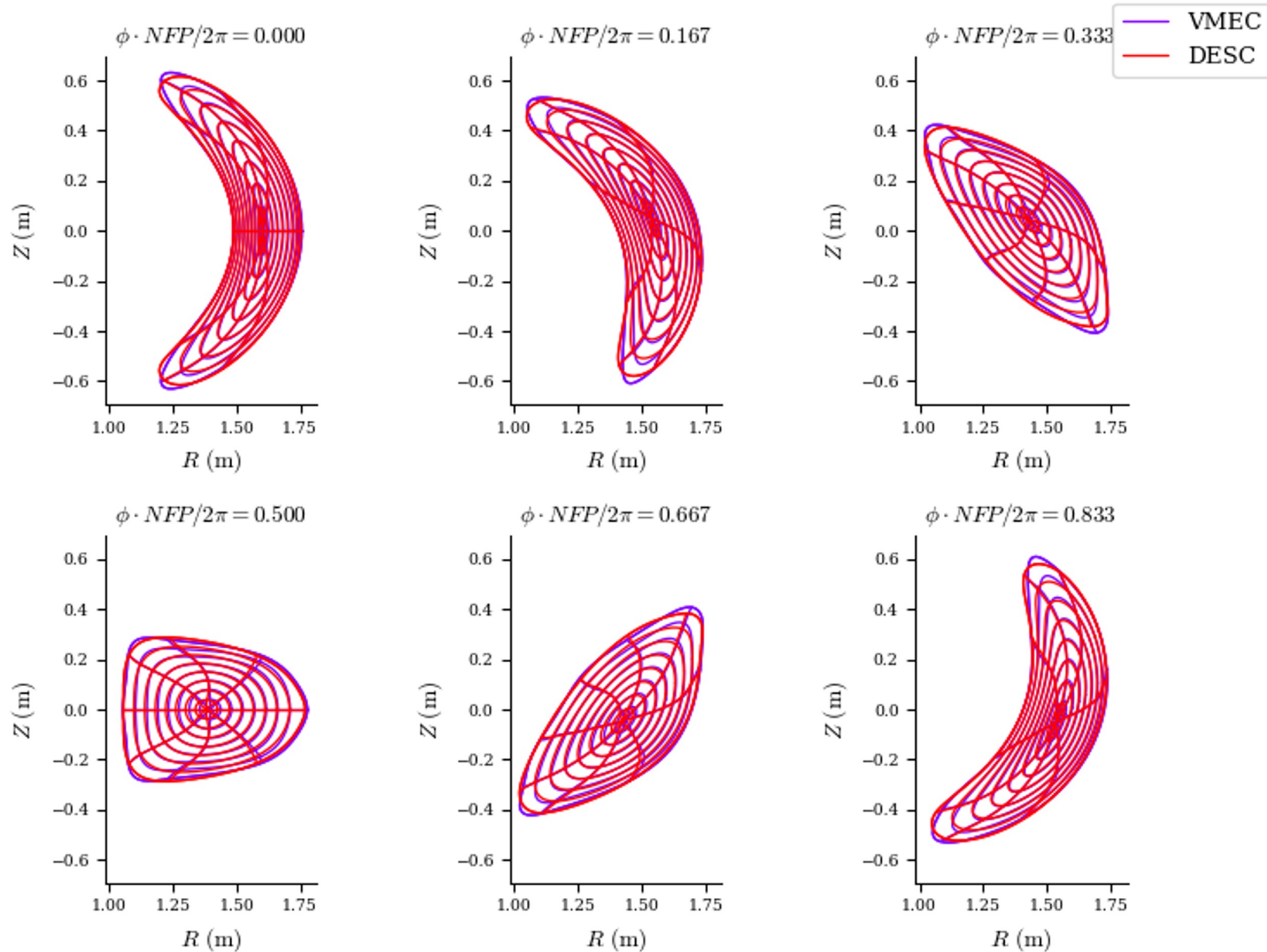
$B_{\text{in}} = B$ from fixed boundary DESC

Virtual Casing: $-nxB_{\text{out}} = K_{\text{out}}$

Biot-Savart(K_{out})= B_{coil}

- Alternatively:
- Use high order singular integration scheme (Malhotra et al 2019) to compute virtual casing for B_{out}
- Just check for pressure balance.

Free Surface DESC vs VMEC

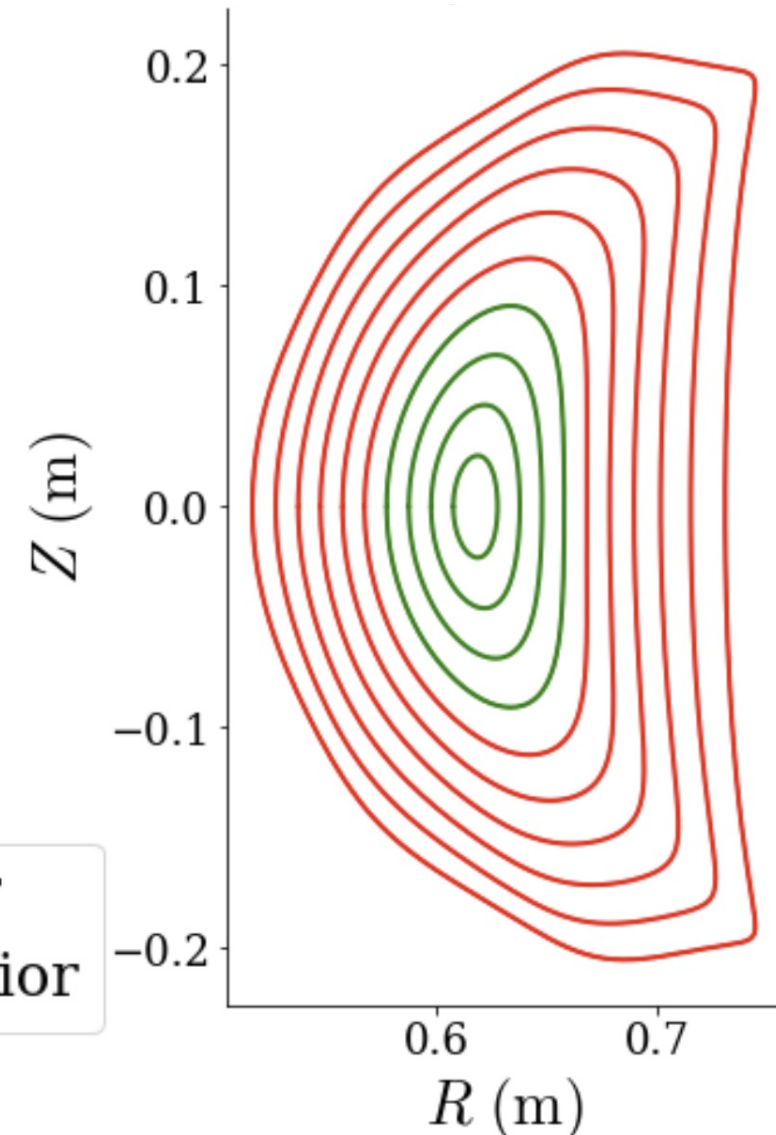


Near-Axis Expansion (NAE) Constraints in DESC

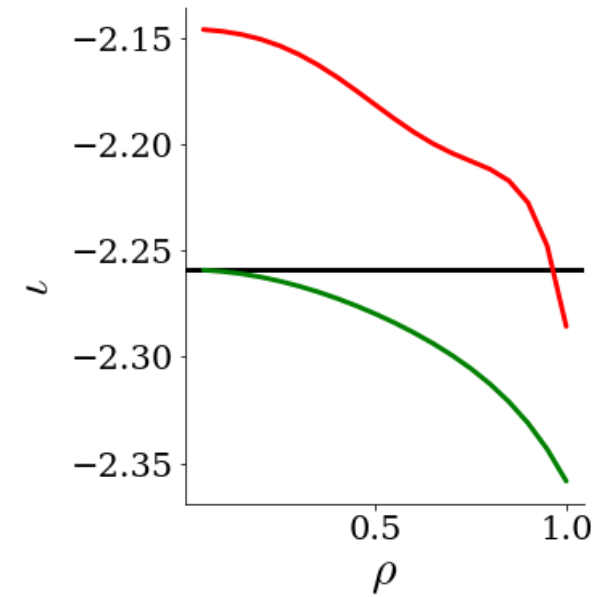
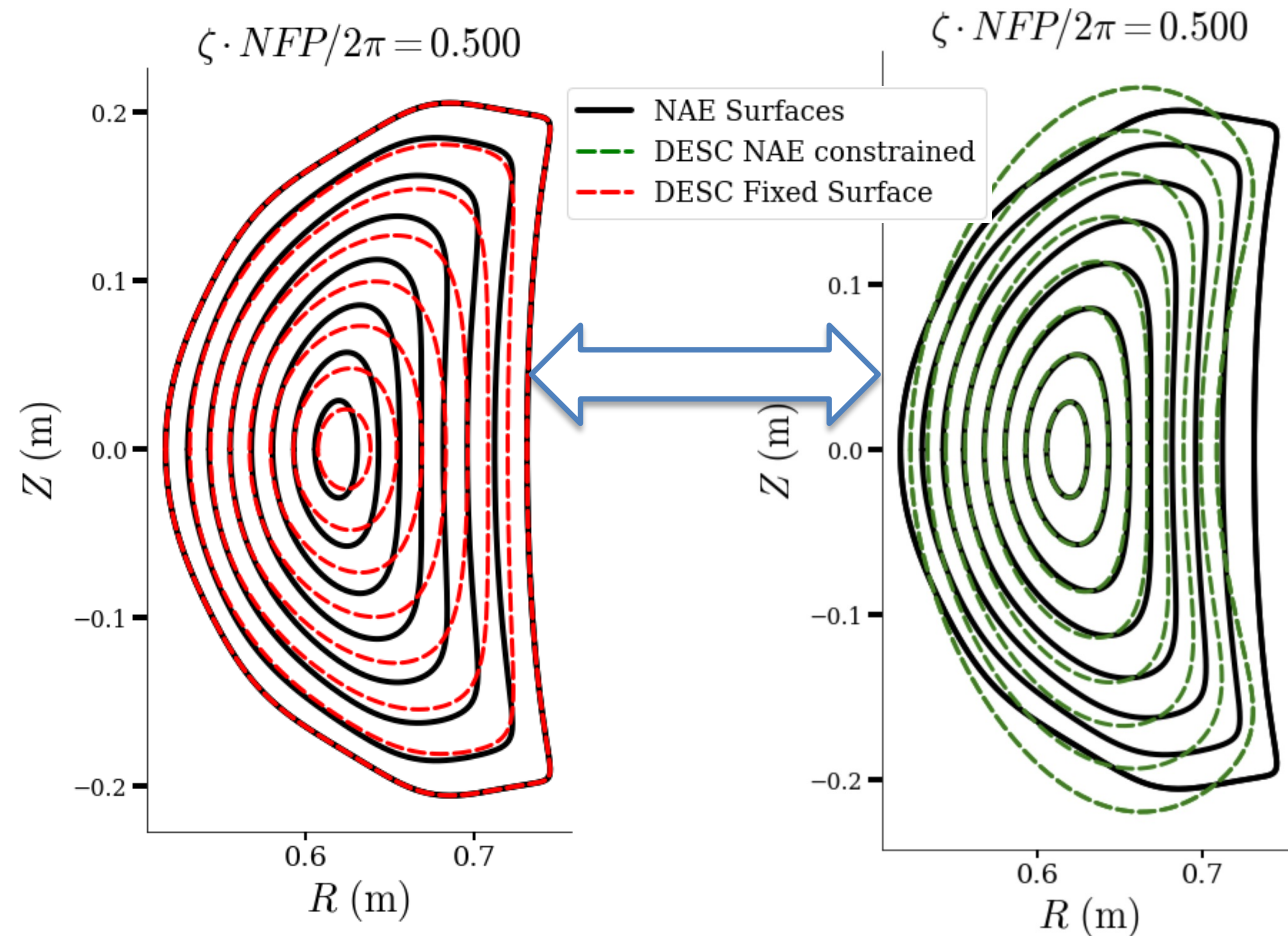
- **Constrain global equilibrium by NAE behavior as $\rho \rightarrow 0$**
 - Use information from NAE where it is **most valid**
 - Avoid singular behavior present when evaluating at **large r**
- **Map NAE coefficients to Fourier-Zernike modes of DESC to fix $O(\rho^0)$ (axis) and $O(\rho^1)$ behavior**



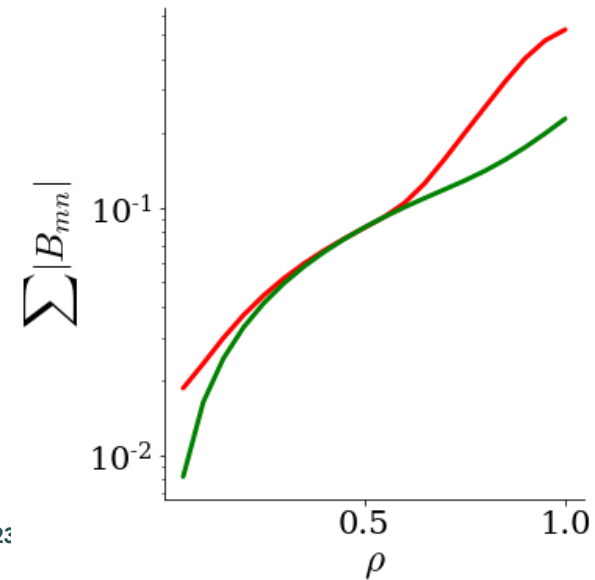
NAE equilibrium evaluated at $r = 0.1$



Near-Axis-Expansion Constrained Equilibria in DESC



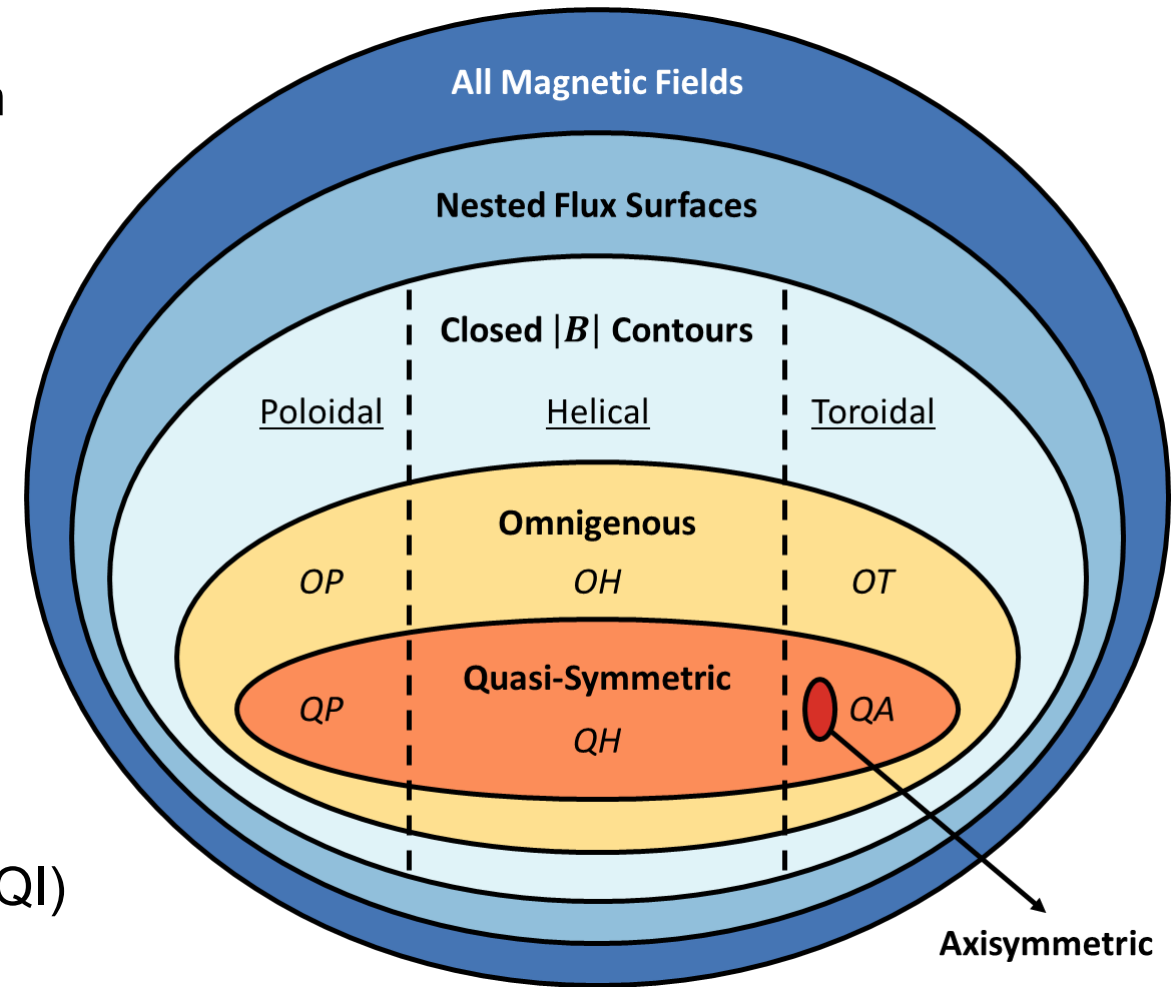
- **NAE-constrained equilibrium iota matches NAE near axis**



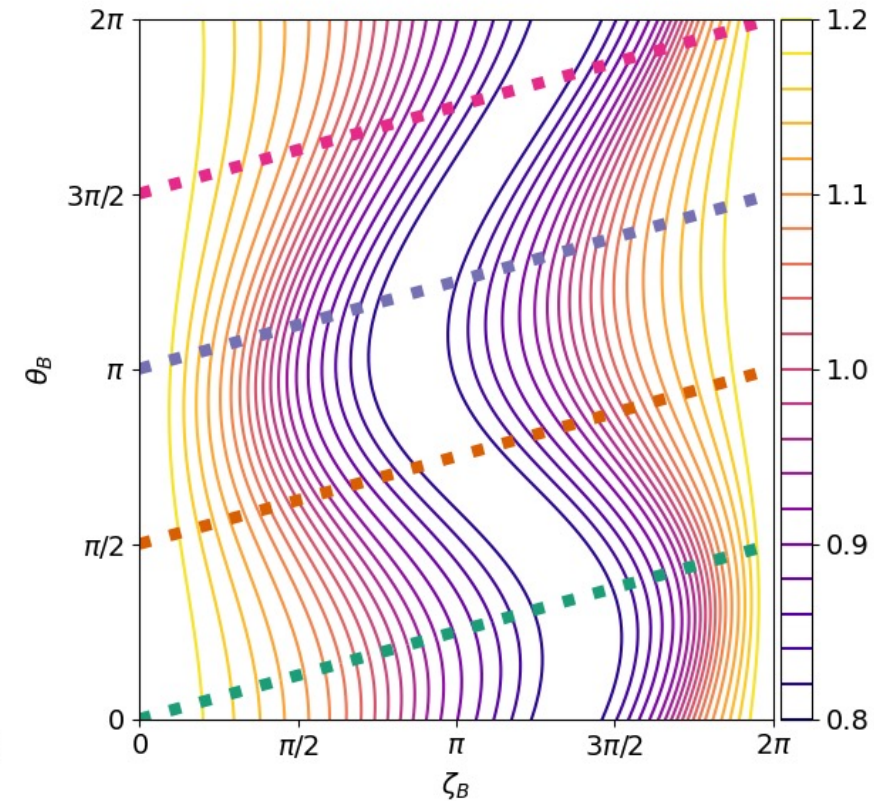
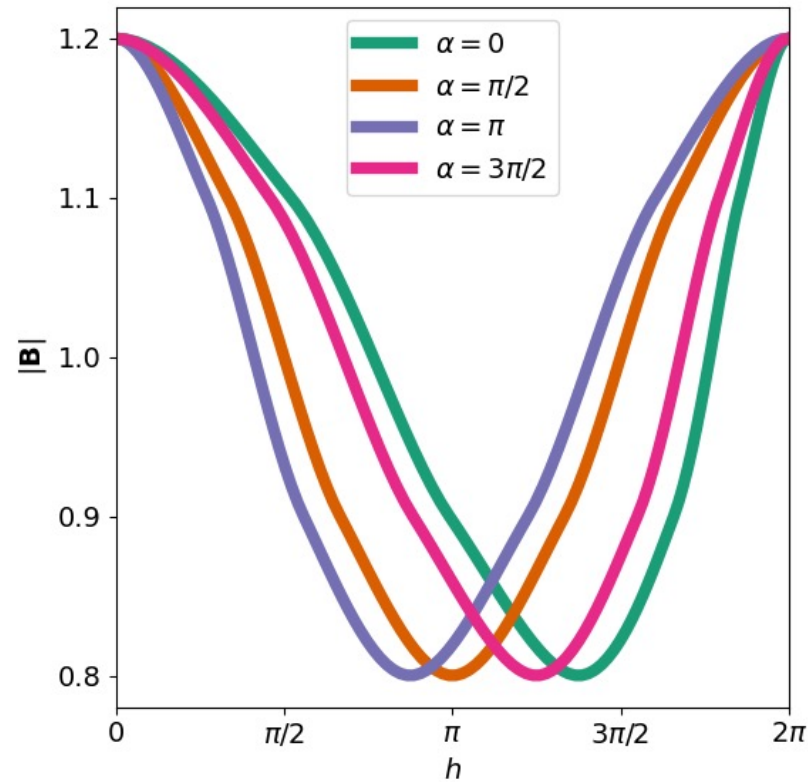
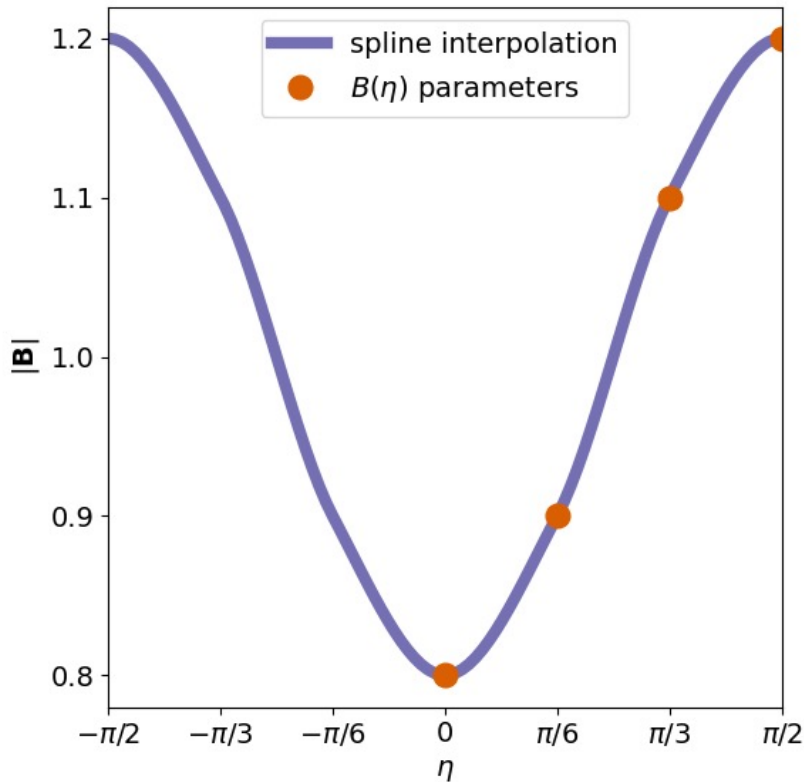
- **NAE-constrained equilibrium maintains better QS near axis**

General Omnigenity: confined particles without Quasi-Symmetry

- **Omnigenity:** the class of magnetic fields in which the bounce-averaged radial drifts of trapped particles vanish
 - $B = |\mathbf{B}|$ contours are closed curves
 - B_{max} contour is straight in Boozer coords
 - “Bounce distances” between consecutive B points are independent of the field line α
- Previous omnigenous equilibria were limited to:
 - Quasi-Axisymmetry (QA)
 - Quasi-Helical symmetry (QH)
 - Omnigenity with Poloidal contours (OP aka QI)
- General omnigenity = larger design space!



General omnigenity optimization is implemented in DESC



1. Define a target magnetic well “shape” (in computational coordinate η)
2. Define a target “shift” on each field line (preserves constant bounce distances)
3. Optimize to minimize the errors: $B_{equilibrium} - B_{target}$

Omnigenity is parametrized by a coordinate mapping

- Map between computational and Boozer coordinates:

helicity = (M, N)

$$h(\rho, \alpha, \eta) = h(\theta_B, \zeta_B)$$

$$2\eta + \pi + \sum_{l=0}^{L_\rho} \sum_{m=0}^{M_\eta} \sum_{n=-N_\alpha}^{N_\alpha} x_{lmn} T_l(2\rho - 1) \mathcal{F}_m(\eta) \mathcal{F}_{nN_{FP}}(\alpha) = \begin{cases} N\zeta_B & \text{for } M = 0 \\ -\theta_B + \frac{N}{M}\zeta_B & \text{for } M \neq 0 \end{cases}$$

ρ = flux surface label
 α = field line label
 η = coord along field line

free parameters (points to x_{lmn})
 Chebyshev polynomials (points to $T_l(2\rho - 1)$)
 Fourier series (points to $\mathcal{F}_m(\eta)$ and $\mathcal{F}_{nN_{FP}}(\alpha)$)

- Constant bounce distances are preserved:

$$\delta \propto \Delta h = h(\rho, +\eta, \alpha) - h(\rho, -\eta, \alpha)$$

$$= 4\eta + \sum_{l=0}^{L_\rho} \sum_{m=0}^{M_\eta} \sum_{n=-N_\alpha}^{N_\alpha} x_{lmn} \left[T_l(2\rho - 1) \mathcal{F}_{nN_{FP}}(\alpha) \underbrace{[\mathcal{F}_m(+\eta) - \mathcal{F}_m(-\eta)]}_{= 0 \text{ because } \sum_{m \geq 0} \mathcal{F}_m(\eta) \text{ is an even function of } \eta} \right]$$

$$= 4\eta$$

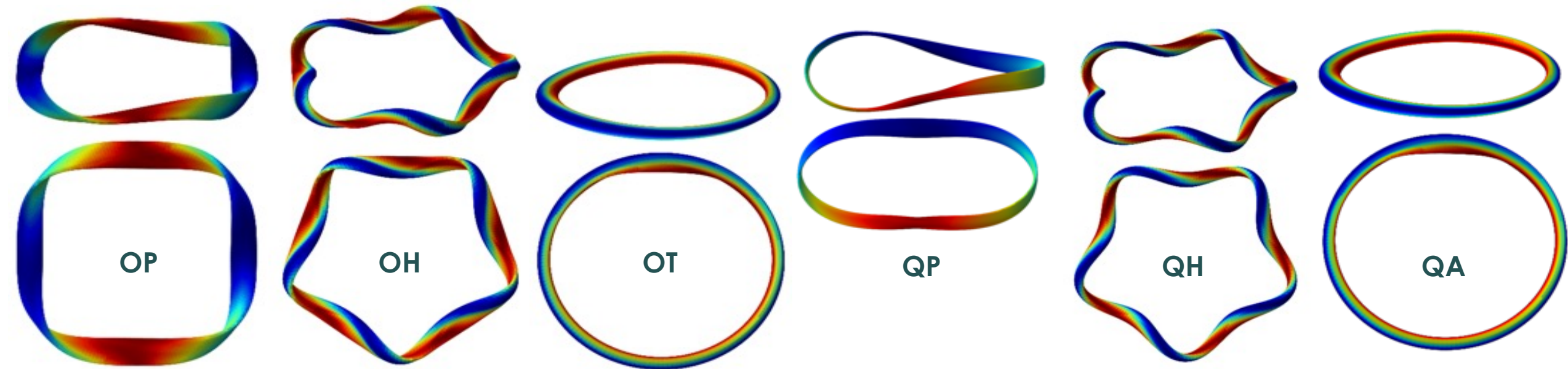
$$\therefore \frac{\partial \delta}{\partial \alpha} = 0$$

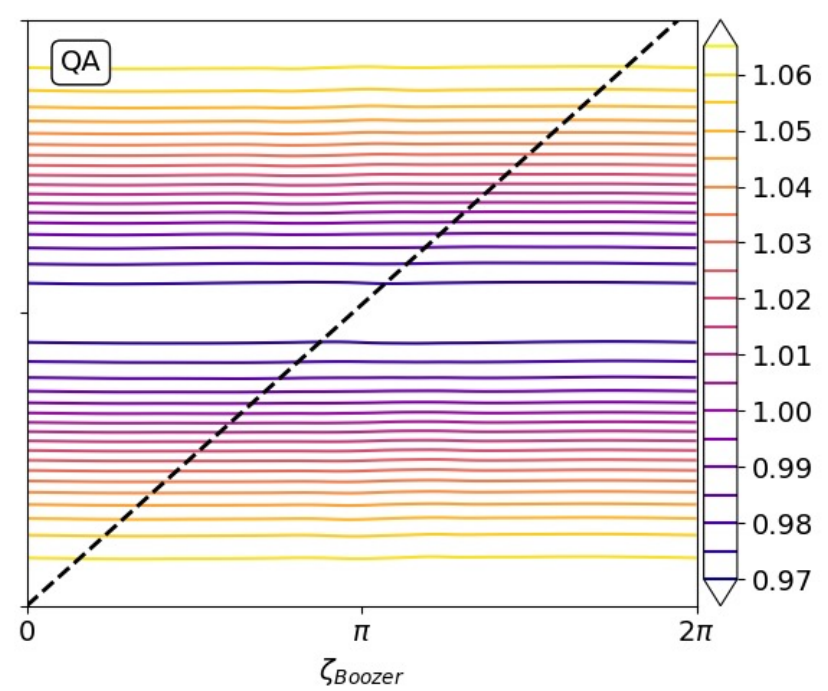
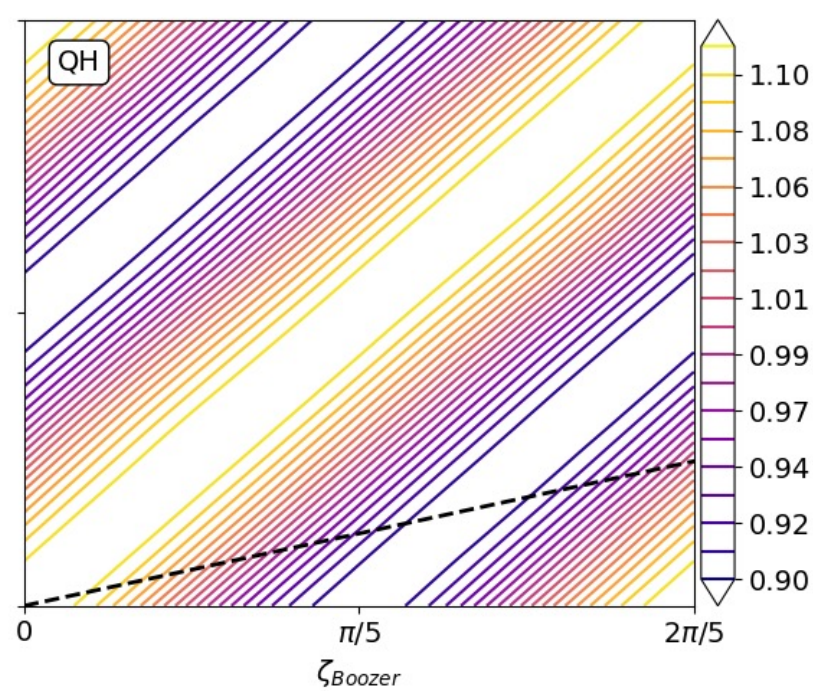
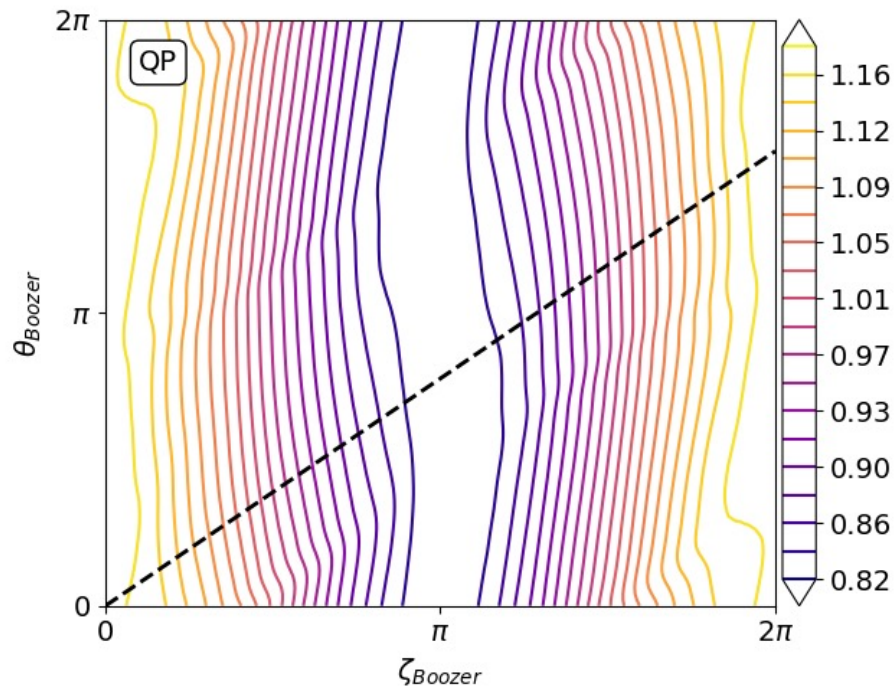
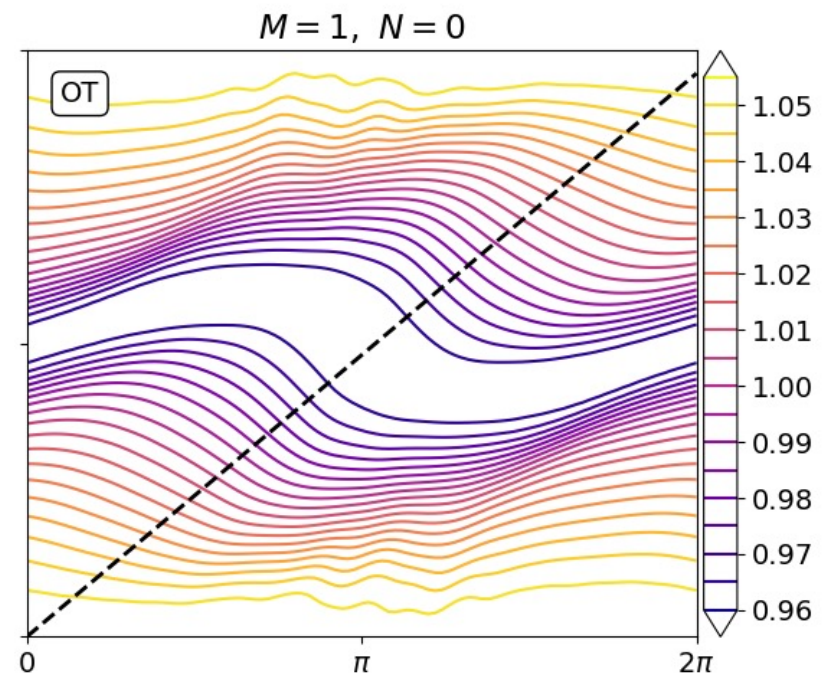
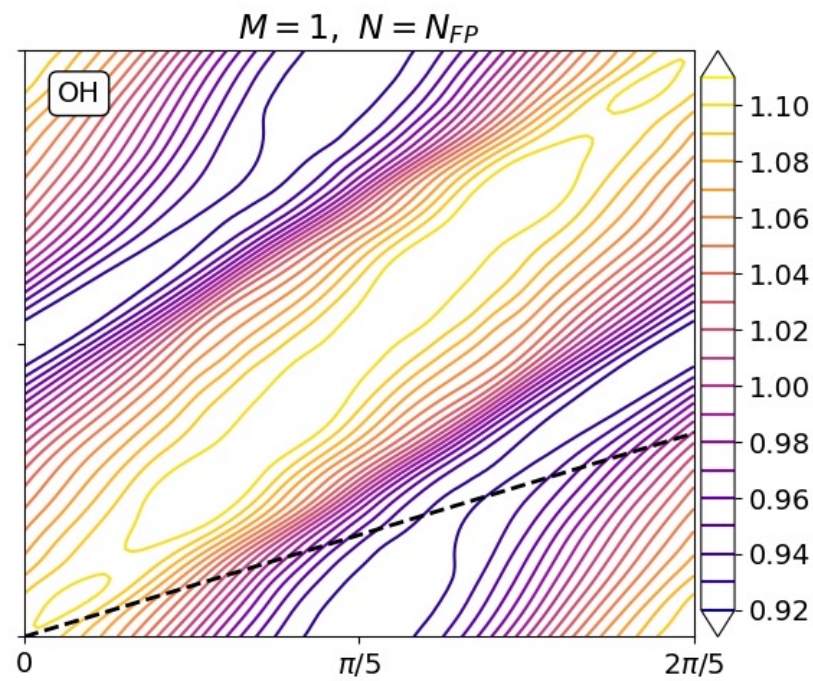
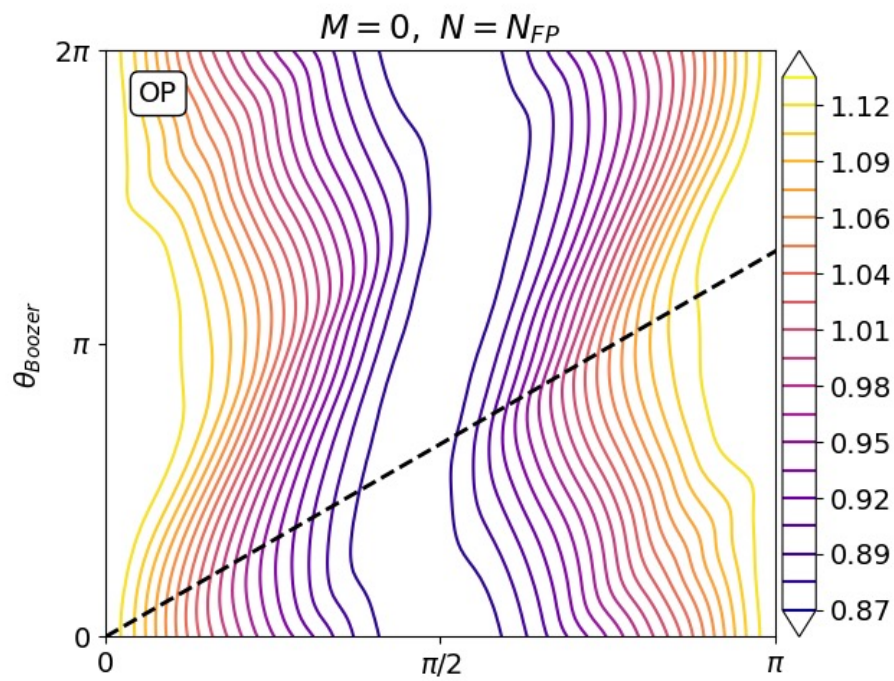
Examples of the different types of omnigenity

General Omnigenity with poloidally/helically/toroidally-closed B contours (OP, OH, OT)

Quasi-Symmetry is a subspace of omnigenity (QP, QH, QA)

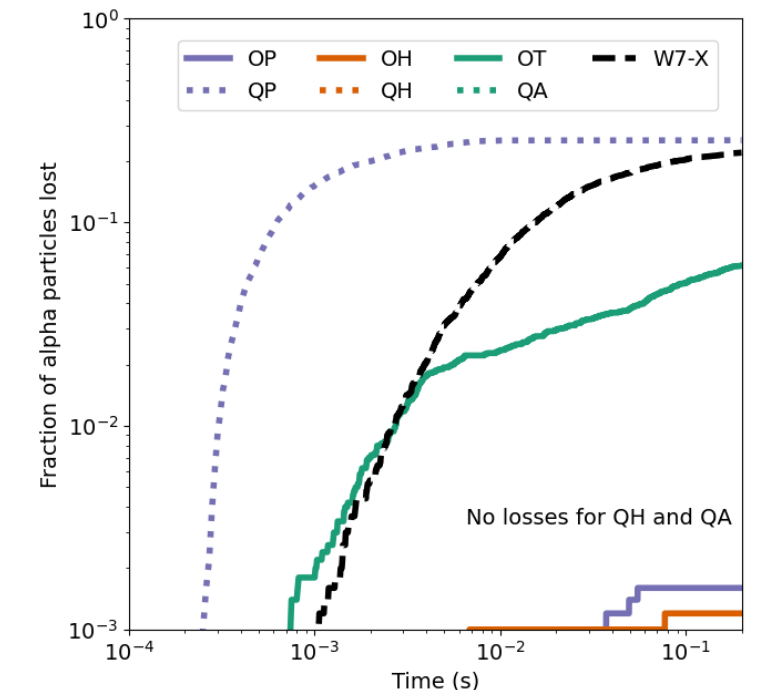
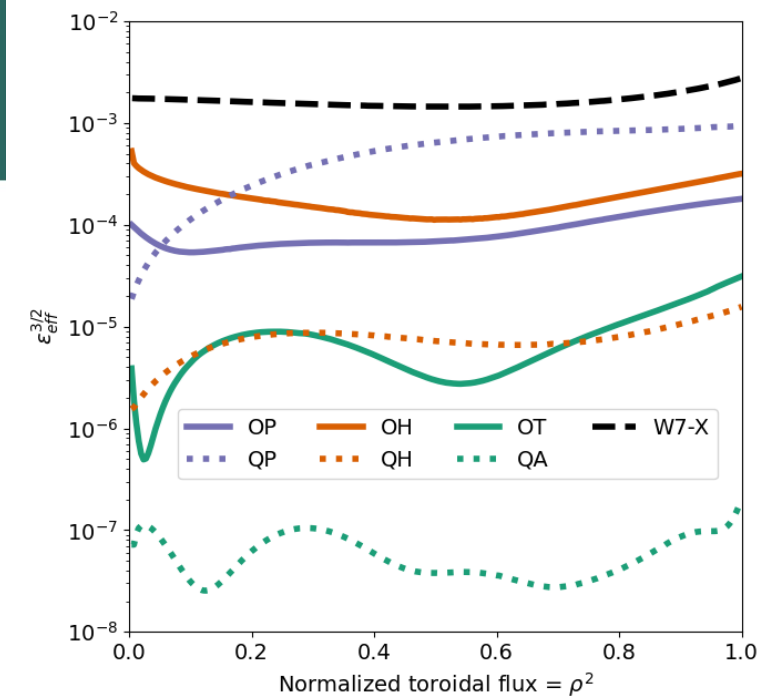
- All solutions used near-axis expansions as an initial guess (from pyQIC and pyQSC)
- Optimized for omnigenity at five surfaces (no other objectives)
- Aspect ratios ≈ 20





Solutions show good particle confinement!

- **Top:** Neoclassical collisional transport magnitude
 - Computed by NEO
- **Bottom:** Collisionless losses of fusion-born alpha particles
 - Computed by SIMPLE
 - Particles initialized at $\rho = 0.5$
 - Configurations scaled to a and B of ARIES-CS
- Reference case is W7-X at $\beta = 4\%$
- Precise QP is difficult (impossible?) to achieve
- Higher alpha particle losses for OT case might be due to wide banana orbits?



GPU Allows Direct Optimization of Particles (Instead of proxies such as omnigenity)

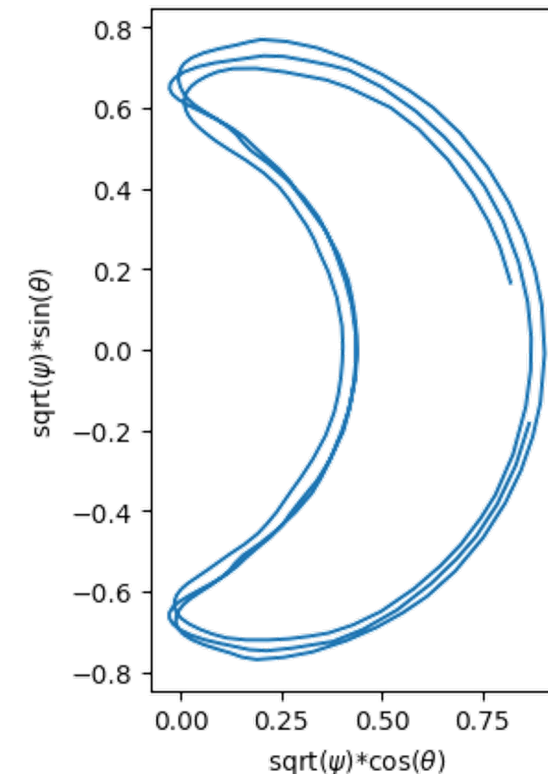
- GPU has advantage in doing the same compute many times
- *Integrate lots of particle trajectories in DESC (guiding center equations of motion)*
- JAX autodiff: Jacobian of trace particle trajectories wrt equilibrium

$$\dot{\psi} = \frac{m}{qB^3} \left(\frac{v_{\perp}^2}{2} + v_{\parallel}^2 \right) \mathbf{B} \times \nabla B \cdot \nabla \psi$$

$$\dot{\theta} = \frac{v_{\parallel}}{B} \mathbf{B} \cdot \nabla \theta - \frac{m}{qB^3} \left(\frac{v_{\perp}^2}{2} + v_{\parallel}^2 \right) \mathbf{B} \times \nabla B \cdot \nabla \theta$$

$$\dot{\zeta} = \frac{v_{\parallel}}{B} \mathbf{B} \cdot \nabla \zeta$$

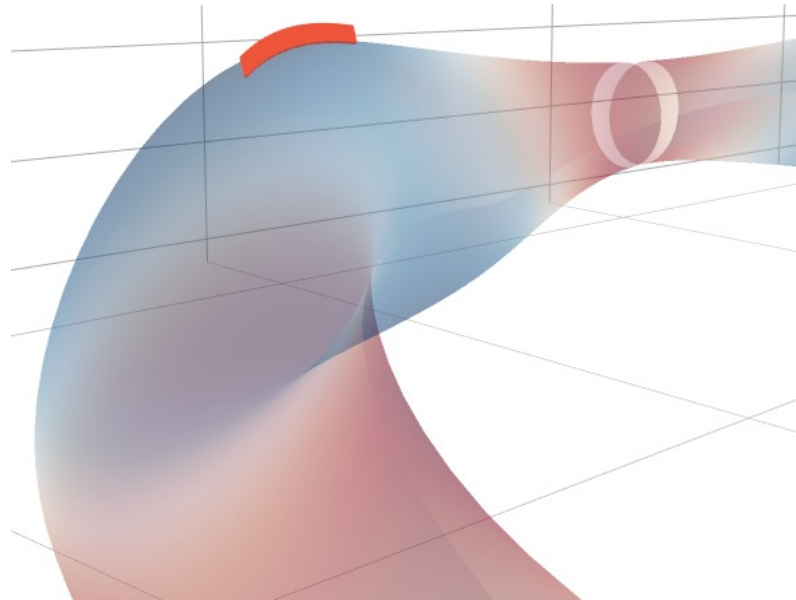
$$\dot{v}_{\parallel} = -\frac{v_{\perp}^2}{2B} \left(\frac{\mathbf{B}}{B} + \frac{mv_{\perp}^2}{2qB^3} \frac{1}{v_{\parallel}} \mathbf{B} \times \nabla B \right) \cdot \nabla B$$



Equilibrium optimization using a Particle Tracer

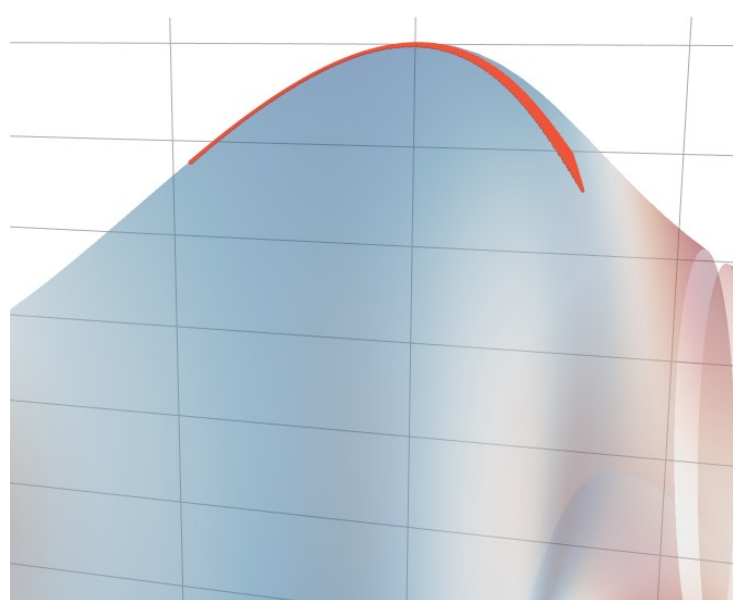
- Particle Tracer:
 - Integrate Guiding Center EoM directly
 - Optimize the equilibrium from particle's trajectories using JAX autodiff

Starting Equilibrium

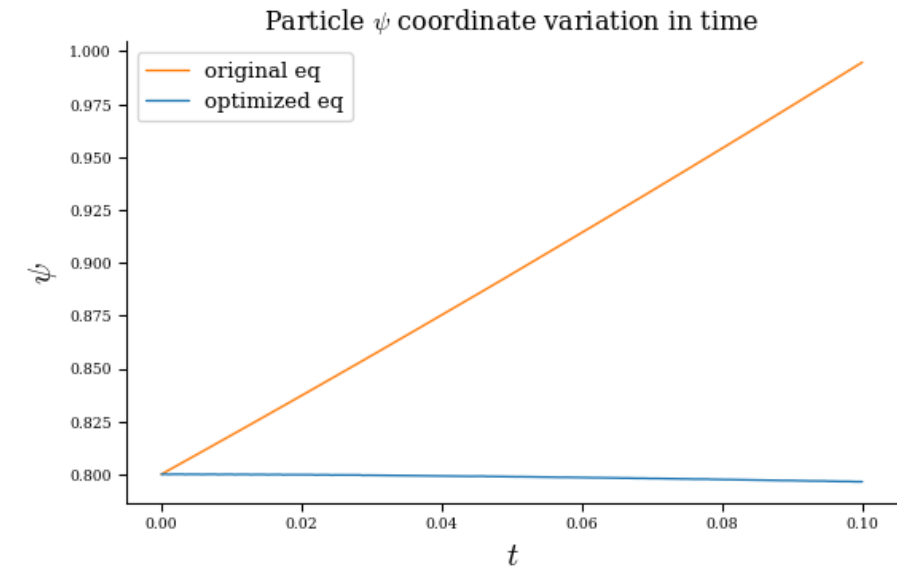


Particle drifting from flux surface ψ

Optimized Equilibrium

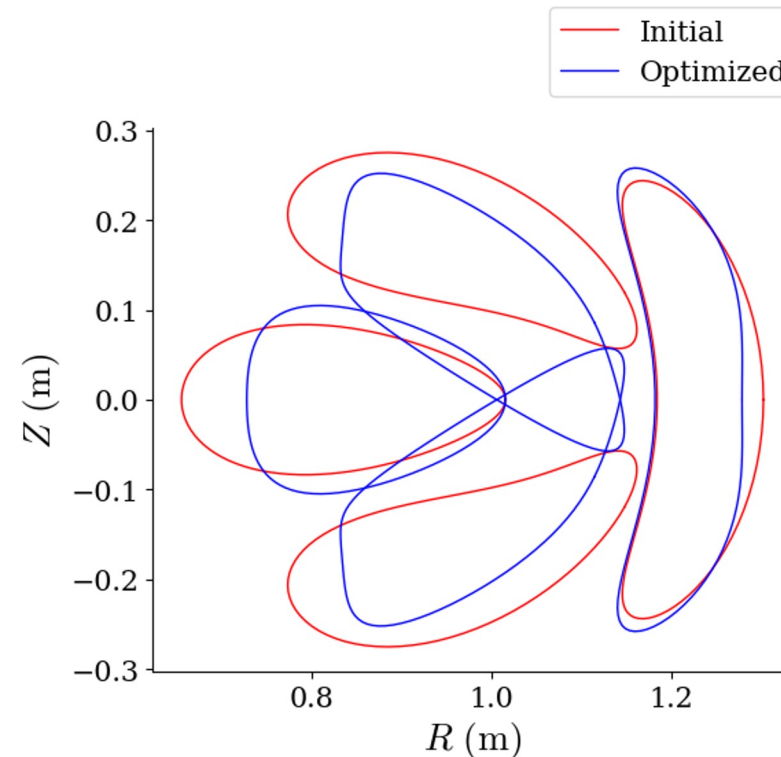
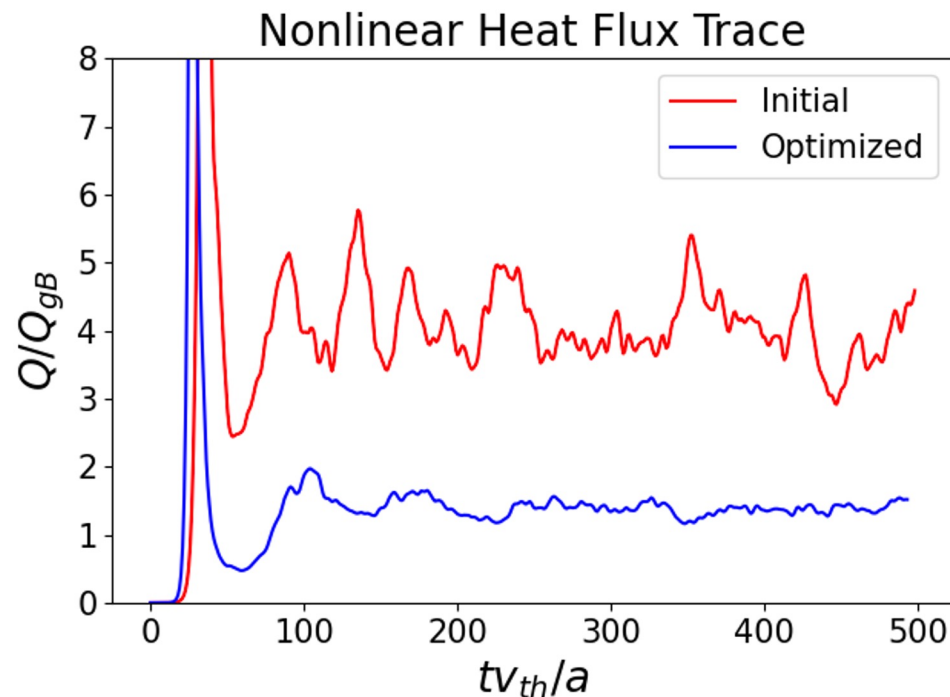


Particle confined in flux surface ψ

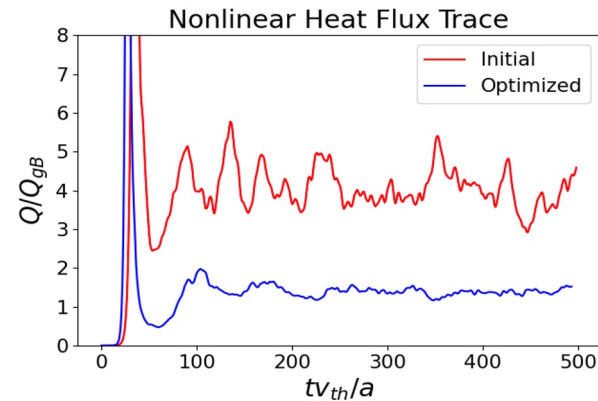
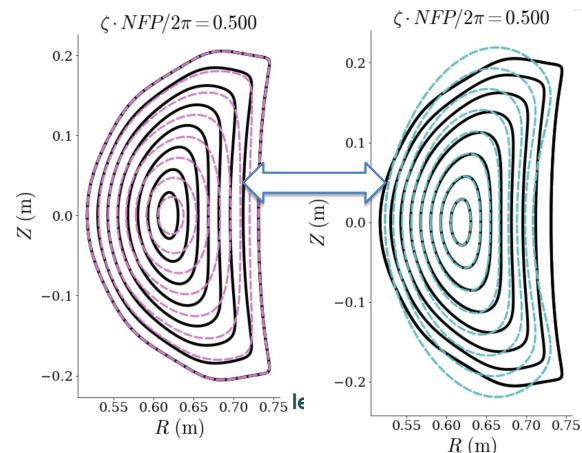
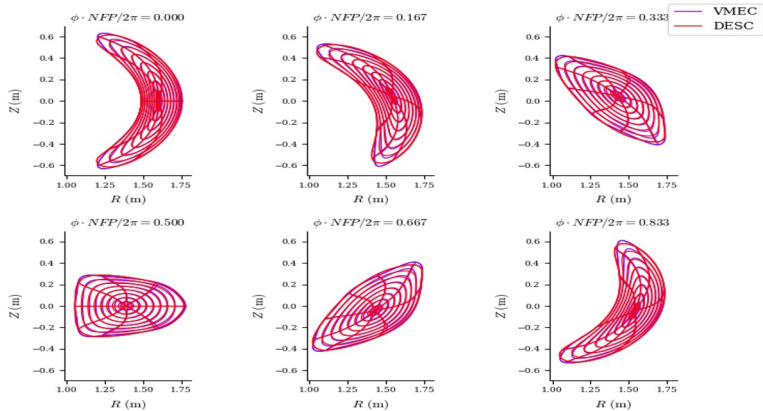
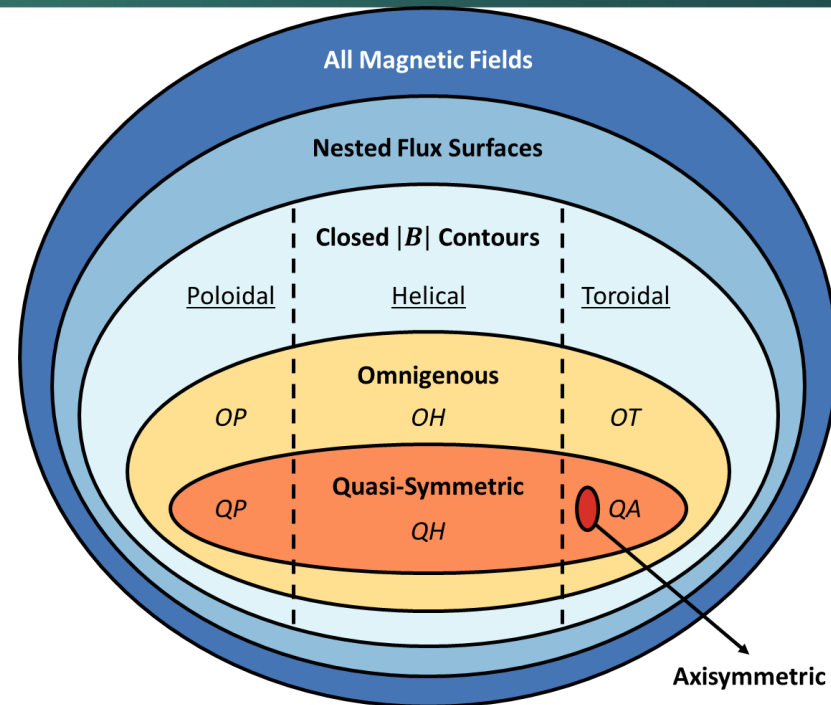
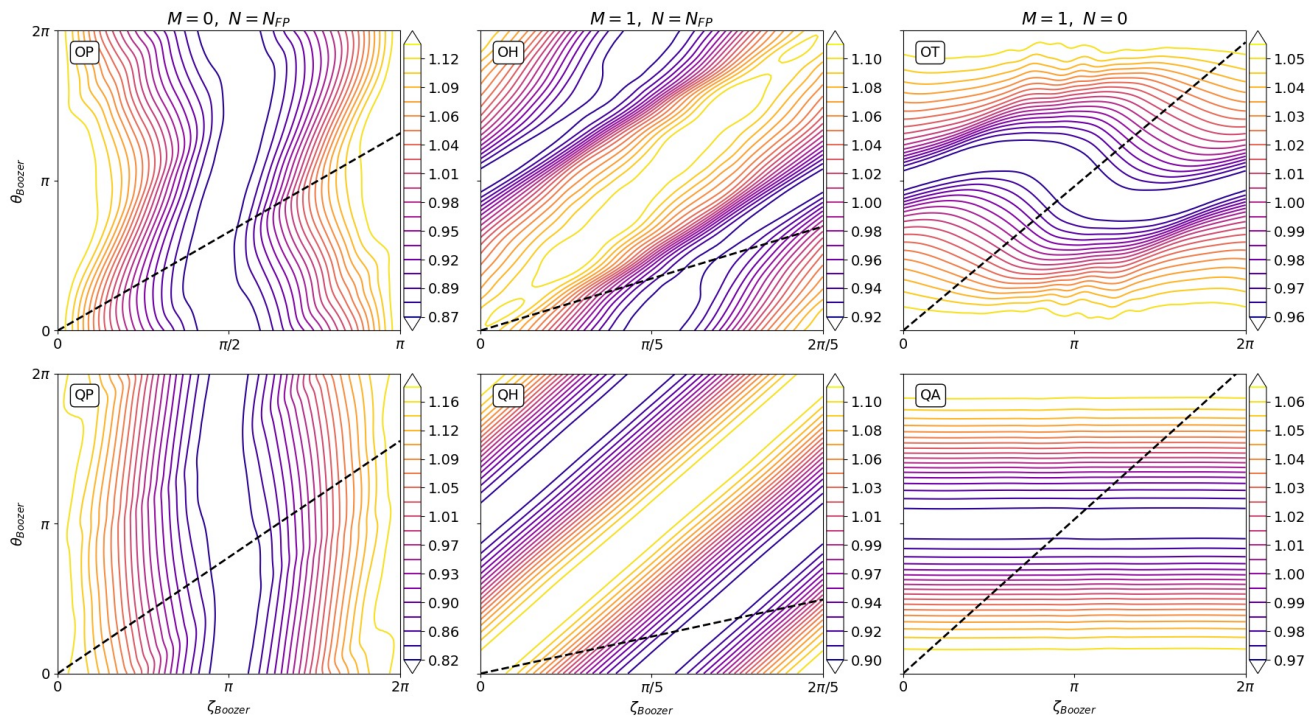


Turbulence + QS Optimization (Using GX+ DESC)

- **GX + DESC coupling enables direct optimization of nonlinear heat fluxes with good quasi-symmetry.**
- **SPSA algorithm allows for cheaper gradient approximations for noisy objectives.**
- **Optimizer reduces nonlinear heat flux by a factor of 3, while maintaining good quasi-symmetry.**



Conclusions: DESC multi-objective optimization applied to Turbulence Optimization, NEA, free surface stellarator & General Omnigenity



Postdoc Opportunity!

- **Novel Stellarator Designs**
- **Stellarator/Divertor Topology**
- **Plasma SOL**
- **Plasma, Applied Math, CS welcome**
- **Work on DESC**

- **Email: ekolemen@princeton.edu (or talk to me)**